## Copper-rich alteration of micrometeorites collected near Calama, Atacama Desert

FIEN MARIE RAF JONCKHEERE<sup>1</sup>, LISA KRÄMER RUGGIU<sup>1</sup>, JÉRÔME GATTACCECA<sup>2</sup> AND STEVEN GODERIS<sup>1</sup>

<sup>1</sup>Vrije Universiteit Brussel <sup>2</sup>CNRS, Aix-Marseille Univ, IRD, INRAE, CEREGE

This study investigates terrestrial alteration of micrometeorites using copper concentrations as a tracer. We analysed micrometeorites from the Atacama Desert near Calama, Chile, 25 km from the Chuquicamata mine, world's largest open pit copper mine. Bulk sediment  $\mu$ XRF analysis performed at AMGC revealed a copper-rich composition. To study the micrometeorite flux and preservation in a hot desert, sediment was sampled from a 1 m² surface at 2 cm depth, using 500 g in this study. 500 particles in total were extracted across multiple size fractions between 2 mm and 63  $\mu$ m. The major element composition of these particles is mapped using  $\mu$ XRF and 173 promising spherules were selected for SEM-EDS imaging and analysis, while 9 were selected for further oxygen isotope analysis.

SEM-EDS revealed well-preserved S-type (14), I-type (5), G-type (12) micrometeorites, showing typical micrometeorite texture like magnetite dendrites in a glassy matrix (G-type), metal beads (I-type) and olivine crystals inside glassy matrix (S-type). These spherules show bulk copper content up to 14 wt% (2.79 wt% average), with Cu-bearing phases observed as rounded inclusions on polished sections ranging from the nm-scale up to a few μm, or as homogeneous replacement of single-phase spherules. The copper-minerals range from covellite (CuS), bornite (Cu<sub>5</sub>FeS<sub>4</sub>), and chalcopyrite (CuFeS<sub>2</sub>), with minor contributions of As and Sb, which has not been described in literature as a typical weathering feature of micrometeorites.

Triple-oxygen isotope analysis using SIMS (Figure 1) assessed parent body origin and atmospheric entry effects. The  $\Delta^{17}O$  vs d<sup>18</sup>O plot [1] shows spherules of major micrometeorite groups. Most particles link to carbonaceous chondrites, and one I-type falls between the TFL and ordinary chondrite fields. Five spherules (without copper) plot near the TFL ( $\sim \delta^{17}O = 0.52 \text{ x} \delta^{18}O$  [2]), suggesting a terrestrial overprint. Remarkably, five Cu-bearing spherules preserve partial extraterrestrial oxygen signatures. Their textures combined with their major composition and isotopic data suggest that those spherules could have an extraterrestrial origin while showing Cu-rich terrestrial weathering.

**References:** [1] Suavet et al. (2010) Earth Planet. Sci. Lett. 293:313-320. [2] Clayton (1993) Annu. Rev. Earth. Planet. Sci. 21:115-149. [3] Thiemens and Brenninkmeijer (1995) Geophys. Res. Lett. 22(3):255-257.

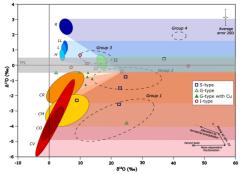


Figure 1: Triple oxygen data by SIMS of the micrometeorites plotted with  $\delta^{16}O(\%s)$  versus  $\Delta^{17}O$  (%s). S., G., and 1-type micrometeorites are plotted in the figure in relation to the micrometeorite groups and meteorite parent body sources. The solid line and grey rectangle represent the TFL  $(\sim 5^{12}O = 0.52 \times 3^{10}O)$ , [21). The attra represents the present average atmospheric incorpic composition of oxygen around the stratosphere-mesosphere transition ( $\delta^{16}O = 25.5\%s$  and  $\delta^{17}O = 11.3\%s$ , [31). The plot is adapted from [11]. The average error is 2SD.