Deciphering mantle vs magma degassing in 3 Ga-old lunar basalts using halogen (F, Cl, Br and I) contents and Cl isotopes

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Phosphates are relatively common in lunar basaltic rocks and their halogen contents and Cl stable isotopes (expressed as d³⁷Cl: $d^{37}Cl = [(^{37}Cl/^{35}Cl)_{Sample}/(^{37}Cl/^{35}Cl)_{SMOC} - 1]*1000)$ have been the focus of many investigations aiming to assess the volatile inventory in the lunar interior [1,2]. High d³⁷Cl values determined in lunar phosphates have been interpreted in terms of mantle source composition as the contribution in the chemical characteristics of lunar basalts of a KREEP component with particularly high d³⁷Cl values [3, 4]. Nevertheless, there are no d³⁷Cl values measured in the 3 Ga-old lunar basalts which are potentially not contaminated by KREEP material. Therefore, the hypothesis of a KREEP component with high d³⁷Cl values involved in the mantle sources of the lunar basalts is not fully established. Furthermore, other processes including, magmatic or impact-related degassing could produce high d³⁷Cl values [5, 6, 7]. To tackle this issue, we have determined the halogen (F, Cl, Br, I) contents and d³⁷Cl in phosphates from the 3 Ga-old lunar meteorite Northwest Africa (NWA) 4734. Our new data determined in situ by SIMS show a first order trend (decreasing I/Cl an Br/Cl with increasing d³⁷Cl) interpreted as evidence of heavy halogen (Br and I) degassing while apatite was crystallising in a magma body. A second order trend with increasing F/Cl correlated with increasing d³⁷Cl, is observed between the data field of the depleted low-µ basalts and the data field of the NWA 4734 apatites. Such a trend cannot be explained by a mixing between a KREEP-like mantle component and a low-µ (depleted-like) mantle component but is rather interpreted as a degassing trend affecting the mantle while exposed to vacuum.

References:

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