

Is lunar dust producing the late Eocene ^3He anomaly, 36 Ma ago? New insights from extraterrestrial ^3He in iridium poor marine sediments.

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Near the Eocene-Oligocene transition, the two largest recent craters (Popigai 100 km diameter; Chesapeake Bay 40 km) formed. This event yields a 2 Ma enduring ^3He anomaly in sedimentary records of the Northern Hemisphere [1]. The origin of the ^3He rich dust particles [1-3], and the global occurrence of this $^3\text{He}_{\text{ET}}$ signature is still debated. Therefore, we produced a new record of the flux of $^3\text{He}_{\text{ET}}$ and ET iridium (Ir) from a deep marine core located on Maud Rise in the Southern Ocean.

The micro-fossil ooze from ODP 689B core contains very low terrigenous mineral content and very low iridium background (5-15 pg/g [4]). These Ir concentrations are close to the expected 9 pg/g Ir_{ET} as calculated using sedimentation rate (0.62 cm/ka), dry density (0.7 g/cm³) and current global/annual Ir_{ET} flux (4 pg Ir cm²/ka). At the CRPG Nancy noble gas laboratory, we measured 23 samples following the standard protocol for helium analysis [1,5]. Decarbonized aliquots were degassed at 1600°C in the ultra-high vacuum induction furnace system [6]. After gas purification, ^3He and ^4He abundances were measured with the home-tuned CRPG Split Flight Tube mass spectrometer [5,7]; data shown in Fig. 1.

The ^3He -flux (Fig. 2) supports that the $^3\text{He}_{\text{ET}}$ -anomaly is a worldwide phenomenon. The ^3He concentrations increase by a factor of 4 whereas the Ir concentrations [4] remain at very low background values never exceeding 20 pg/g Ir. To constrain the $^3\text{He}/\text{Ir}$ ratio during the anomaly, highly siderophile element concentration will be measured in sample aliquots.

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Figure 1: Measured $^3\text{He}/^4\text{He}$ vs ^3He in the non-carbonate fraction of the sediment core IODP 689B.

Figure 2: $^3\text{He}_{\text{ET}}$ flux across the Eocene - Oligocene transition.

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