Reassessment of the age of the solar system and the chronology of chondrule formation

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Chondrules, submillimetre igneous silicate spherules, are the most abundant high-temperature objects formed during the evolution of the circumsolar disk. The chronology of their formation remains however highly controversial. Absolute Pb-Pb ages and ²⁶Al-²⁶Mg ages relative to calcium-aluminiumrich inclusions (CAIs) provide inconsistent chronologies, with Pb-Pb ages showing early and protracted chondrule formation up to 4 Ma after CAIs [1] whereas ²⁶Al-²⁶Mg ages suggest that chondrule production was delayed by at least 1.5 Ma [2-3]. We have developed a new method to precisely determine in situ ²⁶Al-²⁶Mg ages of spinel-bearing chondrules, with spinel being far less sensitive to secondary asteroidal processes compared to the other phases commonly used for ${}^{26}Al - {}^{26}Mg$ dating [2–3]. Our data demonstrate that ²⁶Al-²⁶Mg chondrule formation actually starts 1 Ma earlier than previously thought, i.e., as early as $0.75 \pm$ 0.26 Ma, and extends over the entire lifetime of the disk [4]. Such a shift in chondrule ages relative to CAIs is however not sufficient to reconcile the Pb-Pb and ²⁶Al-²⁶Mg chronologies of chondrule and achondrite formation, as no chondrule yet formed simultaneously with CAIs. Thus, either chondrules' Pb-Pb ages and volcanic achondrites' ²⁶Al-²⁶Mg ages are incorrect or the age of CAIs should be revaluated at 4.568.7 Ma to ensure consistency between chronometers. Given that the canonical age of CAIs was determined using only 4 specimens [5] and older ages up to 4,568.2 Ma have also been measured [6], modelling based on our new spinel-derived ²⁶Al-²⁶Mg ages favours the adoption of 4,568.7 Ma as the new absolute canonical age of CAIs. This enables reconciling both the Pb-Pb, ²⁶Al-²⁶Mg and ¹⁸²Hf-¹⁸²W ages of chondrules and achondrites, highlights a 0.7-1 Ma gap between the formation of refractory inclusions and chondrules, and supports the homogeneous distribution of ²⁶Al in the circumsolar disk.

[1] Bollard et al. (2017) *Sci. Adv.* 3, e1700407. [2] Villeneuve et al. (2009) *Science* 325, 985–988. [3] Fukuda et al. (2022) *Geochim. Cosmochim. Acta* 322, 194–226. [4] Piralla et al. (2023) *Icarus* 394, 115427. [5] Connelly et al. (2012) *Science* 338, 651–655. [6] Bouvier & Wadhwa (2010) *Nat. Geosci.* 3, 637–641.