Archaeal Lipid Hydrogen Isotope Signatures of the Metabolically Flexible *Archaeoglobus fulgidus* During Autotrophy and Heterotrophy

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The hydrogen isotope compositions (δ^2 H) in lipid biomarkers track metabolic processes and can remain stable over geologically relevant time scales, with hydrogen isotope exchange rates ranging from 10^4 to 10^8 years [1]. The lipid δ^2 H in photosynthesizing eukaryotes (e.g., leaf wax δ^2 H) has therefore been widely used for paleoclimate reconstructions [2]. Microbial lipid $\delta^2 H$ may reveal additional information about geobiological processes, as microbes live in a broad range of habitats and have wider metabolic capabilities. While bacterial lipid $\delta^2 H$ has been extensively studied and shown to be quantitatively related to central metabolism [3], archaeal lipid δ^2 H remains largely understudied. In this study, we conducted experiments with pure cultures of an anaerobic. hyperthermophilic, and sulfate-reducing archaeon, Archaeoglobus fulgidus. This strain was grown on different substrates to test the effects of carbon metabolism (autotrophy on CO₂ vs. heterotrophy on lactate) and redox potentials (lactate vs. H_2 as electron donors) on their lipid $\delta^2 H$. All experiments were conducted with three different $\delta^2 H$ values of growth medium to apportion the relative contributions of H from distinct sourcesprotons in water, hydride carrier molecules, and organic substrates—to the final lipid products. Preliminary $\delta^2 H$ data from these experiments indicate that more lipid-bound hydrogens derive from water during autotrophy compared to heterotrophy. In light of the experimental results, we discuss how this approach can include other metabolisms to enable a generalized framework for interpreting archaeal lipid $\delta^2 H$. The empirical framework can be validated by analyzing the $\delta^2 H$ of relevant environmental samples, allowing us to assess the potential application of the H isotope composition of archaeal lipids as paleoenvironmental proxies.

References: [1] Sessions et al. (2004), Geochim. Cosmochim. Acta 86, 1545-1559. [2] Sachse et al. (2012), Annu. Rev. Earth Planet. Sci. 40, 221-249. [3] Wijker et al. (2019), Proc. Natl. Acad. Sci. 116, 12173-12182.