

Exotic Mo isotope composition in the Nuvvuagittuq Greenstone belt

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The isotopes of Mo, a moderately siderophile element (MSE), can potentially be used as a new tracer of mantle heterogeneities produced as a result of late accretion, i.e. the ~0.5% addition of extraterrestrial materials to the Earth's mantle after core formation, which is most commonly invoked for explaining the high contents of highly siderophile elements (HSE) and their chondritic relative abundances in the present-day mantle [1]. The rationale for such an approach is based upon (1) the observation that most extraterrestrial materials have different Mo isotope compositions compared with the present-day accessible Earth's mantle. Therefore, late accretion is expected to have changed the Mo isotope composition of the Earth's mantle. (2) the evidence for the preservation of early mantle heterogeneities, in particular in Archean rocks, such as those from the Nuvvuagittuq Greenstone Belt (NGB) (Canada, ≥ 3.8 Ga) [2,3].

Replicated high-precision Mo isotope data for 6 NGB samples obtained with our new N-TIMS method [4] reveal the first mass-independent Mo isotopic anomalies ever reported in terrestrial rocks. One sample (n=5 replicates) shows 'normal' Mo isotope composition, but also exhibits exceptionally high HSE and Mo contents, which suggests contamination via post-eruption processes or during sample collection and processing. The other five samples show resolvable ⁹⁴Mo and ¹⁰⁰Mo deficits, averaging at -5.0 ± 1.9 ppm, -4.7 ± 3.2 ppm (± 2 SE, n=18), respectively, but no evidence for ⁹⁵Mo and ⁹⁷Mo anomalies ($\mu^{95}\text{Mo} = 0.0 \pm 1.5$, $\mu^{97}\text{Mo} = 0.5 \pm 1.4$). The ⁹⁴Mo and ¹⁰⁰Mo excesses observed in meteorites are coupled with ⁹⁵Mo excesses [5]. Therefore, the exotic Mo isotope pattern of the NGB rocks does not fit with the composition expected for the Earth's mantle before late accretion of meteorites, which should also show significant ⁹⁵Mo deficits in addition to ⁹⁴Mo and ¹⁰⁰Mo deficits.

[1] Walker, *Chemie der Erde*, 2009; [2] O'Neil et al., *Science*, 2008. [3] Touboul et al., *Chem. Geol.*, 2014. [4] Yobregat et al., *JAAS*, in revision; [5] Budde et al., *Nature Astronomy*, 2019.