

Peridotite and pyroxenites from the mantle-wedge underneath the Northern Andes (Mercaderes area, Colombia)

LUCA NOTINI¹, ELISABETTA RAMPONE¹, MARCO SCAMBELLURI², ALBERTO ZANETTI³, FABIO FERRI⁴ AND ANDRÉS RODRÍGUEZ-VARGAS⁵

¹Università di Genova

²Distav

³IGG-CNR, Pavia

⁴Università di Padova

⁵Minerlab, Bogotá

Presenting Author: luca.notini@edu.unige.it

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1. Notini¹, E. Ramponi¹, M. Scambelluri¹, A. Zanetti², F. Ferri^{3,4}, A. Rodríguez-Vargas⁵

¹Dipartimento di Scienze della Terra, dell'Ambiente e della Vita, Università di Genova, Italy

²IGG-CNR Pavia, Italy

³Dipartimento di Geoscienze, Università di Padova, Italy

⁴EIT Raw Materials, South Innovation Hub, Rome, Italy

⁵Minerlab, Bogotá, Colombia

In the Mercaderes – Rio Mayo area of Southern Colombia, the Pleistocenic Granatífera tuff contains peridotite and pyroxenite xenoliths offering a direct view of the supra-subduction mantle underneath the Northern Andes [1-2-3]. The xenoliths are heterogeneous and range from subordinated garnet peridotite to dominant garnet websterite and clinopyroxenite. A few xenoliths display slightly deformed granular structures (evident in garnet pyroxenite) that retain evidence of melt-rock interaction (e.g., anhedral interstitial garnet and pyroxene overgrowing coarse, deformed pyroxene grains). In most peridotite and pyroxenite xenoliths, the coarse granular assemblage is overprinted by porphyroclastic to mylonitic textures, characterized by ortho-, clinopyroxene and garnet porphyroclasts in a mylonitic orthopyroxene + clinopyroxene \pm olivine matrix.

Overall, mineral compositions define large intervals in terms of Mg-value (82-92) and Na contents in clinopyroxene (0.70 - 1.60 wt%). On the other hand, within a single xenolith, the compositions of rock-forming minerals are very homogeneous without core-to-rim or porphyroclast-matrix variations. The Pressure-Temperature estimates from a selected number of peridotite and pyroxenite samples range between 1150°C-1250 °C and 27-35 kbar (two pyroxene and garnet-pyroxene calibrations). The homogeneous porphyroclast and matrix compositions indicate the pervasive re-equilibration of all xenoliths during mylonitic deformation. Overall, the rock textures combined with mineral thermobarometry indicate that the Mercaderes xenoliths derive from a highly deformed mantle-

wedge domain near the lithosphere-asthenosphere boundary, according to a recent geodynamic model of the area [4]. Despite the geodynamic location of the studied xenoliths, no modal metasomatism has been yet observed.

[1] Weber (1998), PhD Thesis. [2] Rodríguez-Vargas (2005), *Lithos* 82(3-4), 471-484. [3] Bloch (2017), *Earth and Planetary Science Letters*, 476, 47-58. [4] Wagner (2017), *Geophysical Research Letters*, 44(13), 6616-6623.