

A study of nitride inclusions in chondrites: Tracers for nebular processes?

JAN LEITNER¹, CHRISTIAN VOLLMER², DENNIS HARRIES³, JÁNOS KODOLÁNYI¹, ULRICH OTT¹ AND PETER HOPPE¹

¹Max Planck Institute for Chemistry

²Institut für Mineralogie, Universität Münster

³Friedrich-Schiller-Universität Jena

Presenting Author: jan.leitner@mpic.de

Nitrogen is the most abundant component in Earth's atmosphere and one of the key elements for the evolution of Earth's biosphere. However, the types of nitrogen carriers, their abundances and N-isotopic compositions are not well constrained for all the potential terrestrial building blocks. Chondritic meteorites contain small quantities of various nitrides. Si₃N₄, TiN, and sinoite (Si₂N₂O) are common accessory phases of enstatite chondrites (ECs). Nitrides in ordinary (OCs) and carbonaceous chondrites (CCs), however, have been scarcely reported [e.g., 1–6]. Here, we report the discovery of Si-, Cr- and Fe-nitride in several carbonaceous chondrites. The average N-isotopic compositions of nitrides from CCs differ significantly from those in ECs [7], and we also observe variations between different CC groups. Generally, nitrogen becomes isotopically heavier in nitrides from EC-(OC-) through CV&CM to CR-CH&CB chondrites. This could be indicative of increasing amounts of outer Solar System nitrogen contributing to the respective Si- and Cr-nitrides, and may reflect different heliocentric distances of the formation regions [8]. We also observe that the N-isotopic compositions of these nitrides are sometimes significantly different from those of their host meteorites and meteorite groups [e.g., 7,9]. This seems to suggest several different nitrogen reservoirs and/or different alteration pathways for the various N-carriers within a given meteorite group, and could also explain the bulk N-isotopic variations observed for, e.g., the OCs [10].

Acknowledgements: The DFG is acknowledged for funding this project in the course of the SPP 1833.

References : [1] Russell S. S. et al. (1995) *Meteoritics* 30, 399. [2] Grady M. M. & Pillinger C. T. (1993) *EPSL* 116, 165. [3] Weisberg M. K. et al. (1988) *EPSL* 91, 19. [4] Weber D. et al. (1994) *Meteoritics* 29, 547. [5] Meibom A. et al. (2007) *ApJ* 656, L33. [6] Harries D. et al. (2015) *Nat. Geosci.* 8, 97. [7] Leitner J. et al. (2018) *GCA* 235, 153. [8] Füri E. & Marty B. (2015) *Nat. Geosci.* 8, 515. [9] Kung C.-C. and Clayton R. N. (1978) *EPSL* 38, 421. [10] Sugiura N. et al. (1998) *Meteorit. Planet. Sci.* 33, 463.