

## **In-situ Lu-Hf and Pu fission track dating of pallasite meteorites**

**THOMAS BURKE, STIJN GLORIE, MARTIN HAND,  
ALEXANDER SIMPSON, SARAH GILBERT AND  
BENJAMIN WADE**

The University of Adelaide

Presenting Author: [thomas.burke@adelaide.edu.au](mailto:thomas.burke@adelaide.edu.au)

Pallasites are stony-iron meteorites that are generally accepted to have originated from the core-mantle boundary of a planetesimal. Few pallasite formation ages have been constrained by whole-rock Re-Os and phosphate Pu fission track dating, which both involve destructive sample preparation and challenging analytical processes (e.g. irradiation at a nuclear facility for fission track dating). Given that only few absolute age determinations exist, it is currently uncertain if the ~130 pallasite meteorites that have been discovered on Earth were sourced from a single planetary body, or multiple planetoids. We have developed a new analytical workflow using  $\mu$ XRF, SEM and LA-ICP-MS/MS that allows in-situ Lu-Hf and Pu fission track dating of phosphate crystals from a polished section of pallasite samples. The main advantages of the approach are: (1) no destructive sample preparation requirements, (2) high spatial resolution and (3) large quantities of data can be collected in a short time span. Fission tracks were analysed using reflective light microscopy and SEM imaging and corrected for contributions from  $^{238}\text{U}$  by in-situ measurement of U concentrations. Cosmic ray tracks were visually distinguished from the fission tracks and discarded to calculate  $^{244}\text{Pu}$  fission track ages. Elevated Yttrium concentrations were detected using  $\mu$ XRF imaging as a proxy for targeting Lu-rich zones in the phosphate grains. Using innovative mass-filtering procedures in an Agilent 8900 reaction-cell mass spectrometer, high order reaction products of  $^{176}\text{Hf}$ ,  $^{175}\text{Lu}$  and  $^{178}\text{Hf}$  were measured free from isobaric interferences.  $^{175}\text{Lu}$  and  $^{178}\text{Hf}$  were measured as proxies for  $^{176}\text{Lu}$  and  $^{177}\text{Hf}$ , respectively. The resulting isotopic ratios of  $^{176}\text{Hf}/^{177}\text{Hf}$  and  $^{176}\text{Lu}/^{177}\text{Hf}$  were used to compute Lu-Hf ages. The technological innovation, involving two radiometric clocks, enables the ability to rapidly obtain age information for a large range of pallasite (and potentially other rocky meteorite) samples, with an aim to enhance our understanding of the early evolution of the solar system.