## Occurrence of detrital monazite within the Esva watershed (NW Spain): EPMA dating and REE geochemistry

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The Navia-Alto Sil (NASD) domain, an under-explored area, is a part of the so-called West-Asturian Leonese Zone (WALZ), which is located in the NW of the Iberian Massif (west-end outcrop of the European Variscides). This wide area ( $^{3}$ ,500 km<sup>2</sup>) is mainly drained by three watercourses (from west to east): Eo, Navia and Esva. The presence of detrital monazite has been identified in the three river mouths, particularly in that of the Esva, whose drainage basin extends 465 km<sup>2</sup>, where the highest monazite proportions in coastal sands were found. Since Rare Earth Elements (REEs) are considered as critical raw materials in relation with emerging technologies, attention has been paid to this finding.

The NASD is geologically constituted by a thick sequence of lower Palaeozoic rocks (mainly sandstones, quartzites and shales with a thin limestone horizon) that lies uncomformably over Precambrian slates. Small-size igneous bodies (granites, diabases, etc) can be found disperse within the metasedimentary record. Detrital monazite appears both as liberated sub-rounded individual grains (100-150  $\mu$ m of diameter) and as mineral inclusions (5-20  $\mu$ m) hosted in metapelitic rock fragments without showing any chemical zoning.

In order to identify the monazite source, some electronprobe micro analysis (EPMA) were carried out over monazite grains from the Esva watershed. Mean  $\Sigma RE_2O_3$  in the monazite is 70.03%; it is enriched in Ce-La-Nd with minor quantities of heavy rare earth elements (from Eu to Lu, including Y), usually below 1% wt. EPMA dating by the (U,Th)/Pb method [1] indicates a Precambrian (Archean) age (~3 Gy), showing the existence of subsequent remobilization processes.

[1] Williams, ML., Jercinovic, MJ and CJ Hetherington, 2007. Microprobe monazite geochronology: Understanding geologic processes by integrating composition and chronology. Annual Review of Earth and Planetary Sciences, 35(1): 137-175.