## <sup>182</sup>W deficits in modern ocean island basalts: vestiges of Earth's accretion?

G.J. ARCHER<sup>1\*</sup>, A. STRACKE<sup>2</sup>, AND T. KLEINE<sup>1</sup>

 <sup>1</sup>Institut für Planetologie, University of Münster, 48149 Münster, Germany (\*archer@uni-muenster.de)
<sup>2</sup>Institut für Mineralogie, University of Münster, 48149

Münster, Germany

Prior studies have reported deficits in <sup>182</sup>W (<sup>182</sup>Hf  $\rightarrow$  <sup>182</sup>W + 2 $\beta$ ; t<sub>1/2</sub> = 8.9 Ma) for some modern ocean island basalts (OIBs) that appear to be negatively correlated with <sup>3</sup>He/<sup>4</sup>He [e.g., 1,2,3]. Proposed explanations for these anomalies have so far included very early differentiation processes (within the first 60 Myr of Solar System history) in Earth's mantle [1] and core-mantle exchange [2,3]. All prior studies reporting <sup>182</sup>W deficits in OIBs relied on TIMS [4], while a prior MC-ICPMS study did not find anomalous <sup>182</sup>W in OIBs [5]. To evaluate whether OIBs have negative <sup>182</sup>W and to further investigate possible causes of these anomalies, we measured the <sup>182</sup>W isotopic compositions of Samoan and Hawaiin OIBs using MC-ICPMS.

The  $\mu^{182}$ W values (ppm deviations of  $^{182}$ W/ $^{184}$ W from terrestrial standards) for Hawaiin and Samoan OIBs measured in our study range from *ca*. 0 to -15, and appear to be roughly correlated with  $^{3}$ He/ $^{4}$ He, consistent with previously reported results [1,3]. Based on these new data, we will discuss models that may explain the  $^{182}$ W- $^{3}$ He/ $^{4}$ He data, including coremantle interaction and the incorporation of mantle material formed within the first 60 Myr of Solar System history. We will also propose that the  $^{182}$ W anomalies may be the result of accessing materials that were accreted early in Earth's history and preserved until the present due to at least partial isolation from the convecting mantle.

[1] Mundl et al. (2017), *Science* **356**, 66-69. [2] Rizo et al. (2019), *Geochem. Persp. Let.* **11**, 6-11. [3] Mundl et al. (2019), *Geochim. Cosmochim. Acta* **271**, 194-211. [4] Archer et al. (2017), *Int. J. Mass. Spectrom.* **414**, 80-86. [5] Willbold et al. (2011) *Nature* **477**, 195.