

The oxidation state of iron in peridotite liquids and implications for planetary magma oceans

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Our knowledge of the relationship between oxygen fugacity (fO_2) and the redox state of iron dissolved in silicate melts comes almost exclusively from investigations of basaltic melts [1 – 3]. Although instructive for understanding mafic magmatism, the mantles of telluric bodies in the inner solar system are broadly peridotitic, not basaltic, in composition [4]. Because the valence of iron may be shifted not only by oxygen fugacity, but also by melt composition [5], understanding the mantle redox states of such bodies necessitates experimental investigation of peridotitic liquids.

In order to wholly fuse peridotite, powders of a synthetic fertile lherzolite composition based on KLB-1 were heated using a gas-mixing aerodynamic levitation furnace to 1900 °C for 30 s under varying Ar+CO₂-H₂ gas mixtures and pure O₂, yielding $\log fO_2$ s between 0 and -8. Here, we report Fe³⁺/Fe²⁺ ratios in the quenched glasses calculated from Fe K-edge XANES spectra collected at the 13-IDE (GSECARS) beamline at APS, Chicago, Illinois.

Calculated Fe³⁺/ΣFe ratios ranged from 0.01 at $\log fO_2 = -8$ to 0.47 at $\log fO_2 = 0$. Dependence of Fe³⁺/Fe²⁺ on $\log fO_2$ conforms to a slope within uncertainty of the ideal value (0.25) and the resultant equilibrium constant for the reaction Fe²⁺O(l) + ¼O₂ = Fe³⁺O_{1.5}(l) is close to unity, indicative of ideal dissolution of both species in peridotite melt. Based on estimates for the FeO contents of planetary mantles [4], bodies in equilibrium with core-forming metal at low pressures should contain 1 – 2 % Fe³⁺ in their mantles. The 3.6 % present in Earth's mantle argues that it underwent additional and/or distinct processes during its accretion.

[1] Kress, V.C., & Carmichael, I.S.E. (1991), *CMP*, **108**, 82-92 [2] Cottrell, E. *et al.* (2009), *Chem. Geol.* **268**, 167-179 [3] Berry, A. J. *et al.* (2018) *EPSL*, **483**, 114-123 [4] Wanke, H. & Dreibus, G., (1988), *Phil.Trans.R.Soc. A.*, **325**, 545-557 [5] Borisov, A.A. *et al.* (2017), *CMP*, **172**, 34