

## Photoelectrochemical generation of perchlorate on Mars: A photostationary state

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The discovery of abundant perchlorate ( $\text{ClO}_4^-$ ) on Mars has resulted in many studies of photochemical  $\text{ClO}_4^-$  generation [1,2]. Some have concluded that it is impossible to explain the abundance of  $\text{ClO}_4^-$  on Mars using generation mechanisms akin to atmospheric gas-phase photochemical processes on Earth [3]. Here, we expand our study of the roles of semiconducting minerals [4] either on the surface or in atmospheric suspension. We show that  $\text{ClO}_4^-$  can be generated via the action of simulated sunlight on anatase and rutile polymorphs of  $\text{TiO}_2$ , both with single crystals and with nanocrystalline films of particles. Rutile, unlike anatase, results in relatively abundant chlorate intermediate, suggesting that the two polymorphs take different reaction pathways. Most importantly,  $\text{ClO}_4^-$  generation from chloride ( $\text{Cl}^-$ ) levels out over time; in experiments with high starting  $\text{ClO}_4^-$ , the concentration decreases to a similar level as when  $\text{ClO}_4^-$  is generated from a zero starting concentration (Fig. 1).

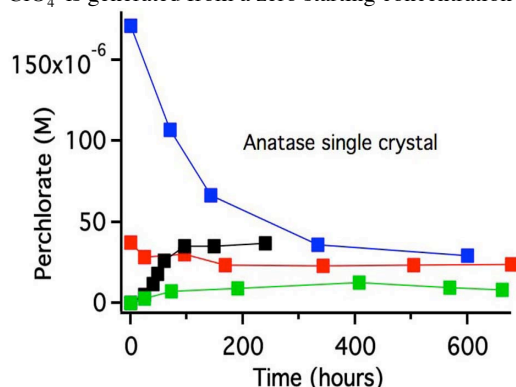


Fig. 1: Production/loss of aqueous  $\text{ClO}_4^-$  as a function of starting  $\text{ClO}_4^-$  concentration. Final  $[\text{ClO}_4^-]$  is similar despite variable initial  $[\text{ClO}_4^-]$ . Black = results from ref. 4, 0.5M  $\text{Cl}^-$ ; Blue, Red, Green = 1.5, 0.7, 2.8 mM starting  $\text{Cl}^-$ , respectively. Analytical error is smaller than symbol size.

The results suggest a photostationary state for perchlorate production; as  $[\text{ClO}_4^-]$  increases the rate of photochemical back-reaction increases until a “steady state” is reached. We have investigated heterogeneous  $\text{ClO}_4^-$  production using hematite as well, and in more Mars-like conditions. If the  $[\text{ClO}_4^-]$  that can be attained photochemically is thus limited, other concentration mechanisms may be required to explain how  $[\text{ClO}_4^-]$  on Mars are similar to those of  $\text{Cl}^-$ .

[1] Hecht et al. (2009) *Science*, **325**, 64–67. [2] Carrier & Kounaves (2015) *Geophys. Res. Lett.*, **42**, 3739–3745. [3] Smith et al. (2014) *Icarus* **231**, 51–64. [4] Schuttlefield et al. (2011) *J. Am. Chem. Soc.*, **133**, 17,521–17,523.