

Total pressure effect on mass-independent fractionation of sulfur isotopes: laboratory experiment of SO₂ photochemistry

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Sulfur mass-independent fractionation (S-MIF) is a useful proxy to understand atmospheric chemistry. Photochemistry of SO₂ accompanies large S-MIF. The isotopic fractionation of SO₂ photochemistry is known to change depending on total pressure [1,2,3], though the detailed mechanism is still uncertain. We examined total pressure dependence of S-MIF by photochemical experiments of SO₂ under various total pressures from 0.01 bar to 1.7 bar. In contrast to previous study [1], our experiment designed for simulating reducing atmosphere including CO and reasonably low pSO₂ condition, two of which are both critical to reproduce S-MIF observed in Archean sediment [4]. The results show that the $\Delta^{33}\text{S}$ and $\Delta^{36}\text{S}$ of photochemical products clearly depend on total pressure. At higher total pressure, the $\Delta^{33}\text{S}$ value is approaching to 0, which is qualitatively consistent with pressure broadening with self-shielding [3]. Also, under the experimental setting, SO produced from the photoexcitation channel is sufficiently smaller than photolysis channel. Thus, the observed total pressure effect probably derived from pressure broadening of SO₂ absorption spectra. Consequently, S-MIF originated from photodissociation channel can be adequately modeled considering two factors: pressure broadening (i.e. total pressure) and self-shielding effect (i.e. pSO₂). Moreover, at high total pressure over 1 bar, the $\Delta^{36}\text{S}/\Delta^{33}\text{S}$ and $\Delta^{33}\text{S}/\delta^{34}\text{S}$ ratios are clearly different from those observed in Archean sediments. The low total pressure of Archean atmosphere is consistent with the previous estimate [5]. Accordingly, S-MIF may provide a useful constraints on total pressure of Archean atmosphere.

- [1] Masterson *et al.* (2011), *Earth Planet. Sci. Lett.* **306**, 253-60. [2] Ono *et al.* (2013), *J. Geophys. Res.* **118**, 2444-54. [3] Lyons *et al.* (2018), *J. Quant. Spectrosc. Radiat. Transf.* **210**, 156-164. [4] Endo *et al.* (2016), *Earth Planet. Sci. Lett.* **453**, 9-22. [5] Som *et al.* (2016), *Nat. Geosci.* **9**, 448-451.