## Differentiated Precursor Bodies as a Source of Earth's Water

J.M. DUDLEY', J.P. GREENWOOD', N. SAKAMOTO', K. ABE'', M. KURODA', J. VAREKAMP', AND H. YURIMOTO'

Dept. Earth & Environmental Sciences, Wesleyan University, Middletown, CT 06459 USA (jdudley01@wesleyan.edu) Natural History Sciences, Hokkaido University, Sapporo, Japan 060-0810

Differentiated planetesimals possibly were Earth's main rocky precursors [1], and Earth's volatiles acquired during accretion [2] may stem from such differentiated materials. Achondrite meteorites provide direct information as to the amount of water accreted before a late veneer. We measured 'H in pyroxene in 5 achondrites to estimate total H<sub>2</sub>O in Vesta, the angrite (APB), and ureilite (UPB) parent bodies. Our results demonstrate that differentiated planetesimals had abundances of water similar to Earth's upper mantle.

The Cameca ims 1270 ion microscope at Hokkaido University with live SCAPS 2-D ion imaging was used to measure 'H in pyroxene in eucrites Juvinas and Pasamonte, volcanic angrite D'Orbigny, plutonic angrite NWA 4590, and ureilite PCA 82506. Spots were pre-sputtered for 30 sec with a 10keV Cs+ primary beam over a 20x20  $\mu m$  area and 'H/ $\!\!$ Si was determined from a central 5x5  $\mu m$  area in 20 cycles. Six pyroxene standards [3] and 1 ppm wt H<sub>2</sub>O glass were mounted in low-temperature metal for calibration. We compare whole-chip mounting of Juvinas (3±3 ppm wt H<sub>2</sub>O), D'Orbigny (10±2 ppm wt H<sub>2</sub>O), NWA 4590 (4±3 ppm wt H<sub>2</sub>O), and PCA 82506 (12±2 ppm wt H<sub>2</sub>O) to grain mounting of Pasamonte (2±3 ppm wt H<sub>2</sub>O), and confirm previous estimates of eucrite and angrite H<sub>2</sub>O [4].

Clinopyroxene-melt partition coefficients, batch melting equations, and literature models of meteorite genesis were used to calculate a range of parent body  $H_{\rm e}O$  contents. Predicted water contents for Vesta (0 to 90 ppm wt  $H_{\rm e}O$ ), the APB (50 to 230 ppm wt  $H_{\rm e}O$ ) and the UPB (200 to 500 ppm wt  $H_{\rm e}O$ ) overlap with estimates of Earth's upper mantle, and support the idea that during the planetesimal-embyro stage of growth, 10's to 100's ppm  $H_{\rm e}O$  could have been accreted by Earth.

[1] Young E. D. (2017) *Nature*, **549**, 461-462. [2] Dauphas N. & Morbidelli A. (2014) http://arxiv.org/abs/1312.1202. [3] Kumamoto K. M. et al. (2017) *Am. Mineral*, **102**, 537-547. [4] Sarafian A. R. et al. (2017) Phil. Trans. R. Soc. *A*, *375*: 20160209.