Early Earth's Atmospheric Evolution

DAVID C. CATLING¹, JOSHUA KRISSANSEN-TOTTON¹, KEVIN J. ZAHNLE²

¹Dept. Earth and Space Sciences/ Astrobiology Program, Univ. of Washington, Box 351310, Seattle WA, USA

²MS-245-3, Space Science Division, NASA Ames, Moffett Field, CA 94035, USA

Atmospheric composition in the Hadean is barely constrained. Zircon data indicate continents, oceans, and maybe life. In models, mean Hadean-Archean climates are not hot because even without land, weathering of seafloor and impact ejecta moderates CO₂ levels and temperature [1, 2].

Rocks indicate oceans, an atmosphere, and land by 3.8 Ga [3], and then, in the Paleoarchean, widespread signs of life. A faint Sun, biospheric gas exchange, and effects of anoxia all bear consideration. Glacial rocks (3.5, 2.9, and 2.7 Ga) suggest global mean temperatures below ~20°C, which can be supported by a CO₂-rich atmosphere (or CO₂-rich with biogenic CH₄). High Archean pCO₂ does not conflict with paleosol inferences with a range 0.02-0.75 bar [4]. Inferred pN₂ <0.5 bar at 2.7 Ga [5] causes a few °C loss of warming, but plausible CO₂ and CH₄ levels can compensate.

A geologic N cycle in an anoxic world may cause low pN_2 . Today, oxidative weathering of organics releases nitrate and nitrate denitrification to N_2 is almost half the long-term N_2 flux; the other half is outgassing [6, p.204]. Archean oxidative weathering was absent, so pN_2 was plausibly lower and would have risen at the Great Oxidation Event (GOE) [5].

What of Archean CH_4 levels? Mass-independently fractionated S isotopes require >20 ppmv [7]. But evidence of H escape to space (light ocean D/H [8] and Xe isotopes that lighten in time [9]), require $2H_2+CH_4$ levels >10³ ppmv.

Finally, global redox is key to the GOE, which requires oxidation of the surface environment by reductant removal. H escape to space driven by CH_4 -enriched Archean air is by far the biggest net oxidation effect over time [10], so its effect of diminishing the pre-GOE O₂ sinks cannot be ignored.

[1] Krissansen-Totton et al. (2018) PNAS, in press. [2] Charnay et al. (2017) EPSL 474, 97-109. [3] Nutman (2006) Elements 2, 223-227. [4] Kanzaki, Murakami (2015) GCA 159, 190-219. [5] Som et al. (2016) Nature Geosc. 9, 448-451. [6] Catling & Kasting, Atmospheric Evolution on Inhabited and Lifeless Worlds, Cambridge Univ. Press, New York, 2017. [7] Zahnle et al. (2006) Geobiol. 4, 271-283. [8] Pope et al. (2012) PNAS 109, 4371-4376. [9] Bekaert et al. (2018) Sci. Adv. 4. [10] Catling et al. (2001) Science 293, 839-843.