

## Asteroid 16 Psyche: Primitive or Differentiated?

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16 Psyche is the target of a NASA Discovery-class mission scheduled to launch in 2022. Interpreting the spectra of M-type asteroids, like Psyche, and reconciling them with the meteorite record has been challenging. They are described as similar to irons, enstatite chondrites (EC) [e.g., 1] and [2] suggested a possible link between the anomalous M-type asteroid 21 Lutetia and CH chondrites. Based on its high radar albedo [3], thermal inertia [4] and estimated grain density [5,6], Psyche is interpreted to be dominated by FeNi and possibly an exposed core of a differentiated asteroid [e.g., 7]. Here we explore an alternative hypothesis that Psyche is a primitive asteroid possibly linked to CH-CB chondrites.

The bulk density of Psyche ( $\sim 4.4 \text{ g/cm}^3$ ) requires considerable porosity ( $\sim 40\%$ ) for interpretation as an iron-rich core [5,8]. Spectra suggest presence of enstatite on its surface [9], as well as a  $3\mu\text{m}$  absorption feature characteristic of hydrated minerals, with similarity to CI chondrites [10]. This led [10] to suggest that Psyche might not be a metallic core or was impacted by a C chondrite-like material. The  $3\mu\text{m}$  feature is present in 35% of M-type asteroids and inconsistent with them as metallic cores [11]. Irons and ECs are not known to contain hydrated minerals. Thus, a link of Psyche, and other M-type asteroids, to irons or ECs seems unlikely. However, presence of magnesian silicates and hydrated minerals is consistent with the metal-rich CH-CB chondrites.

CH-CBs have up to 70 vol.% metal. They contain magnesian silicates and hydrated CI-like clasts [12]. The bulk density of the Bencubbin CB is  $5.66 \text{ g/cm}^3$  [13], closer to that for Psyche than irons. Thus, meteorites like CH-CB are more easily reconciled with the properties attributed to Psyche than irons or ECs and Psyche may be a primitive asteroid.

References: [1] Gaffey and McCord (1977) *Space Sci. Rev.* 21, 555. [2] Moyano-Camero et al. (2016) *Meteorit. Planet. Sci.* 51, 1795. [3] Shepard et al. (2010) *Icarus* 208, 221. [4] Matter et al. (2013) *Icarus* 226, 419. [5] Baer et al. (2011) *Astronom J.* 141, 1. [6] Lupishko (2006) *Solar System Res.* 40, 214. [7] Elkins-Tanton et al. (2015) *Lunar Planet. Sci. Conf.* 46<sup>th</sup>, abstr. 1632. [8] Shepherd et al. (2017) *Icarus* 281, 388. [9] Hardersen et al. [2005] *Icarus* 175, 141. [10] Takir et al. (2017) *Astronom J.* 153, 31. [11] Rivikin et al. (2000) *Icarus* 145, 351. [12] Greshake et al. (2002) *Meteorit. Planet. Sci.* 37, 281. [13] Macke et al. (2011) *Meteorit. Planet. Sci.* 12, 1842.