

Carbon monoxide as a key to understand early Mars and Earth

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Mars' atmospheric CO₂ has been enriched in ¹³C since 4.0 billion years ago. Carbon escape into space may explain the ¹³C-enrichment, though escape via sputtering is not applicable to the early Mars when a magnetic field was present. Alternatively, solar ultraviolet photolysis of CO₂ may cause large carbon isotopic fractionation [1], though the effect had yet to be tested based on laboratory experiment. We have conducted UV photolysis of CO₂ in laboratory and for the first time demonstrate that photolysis of CO₂ by solar-like broadband UV irradiation causes a large negative carbon isotopic fractionation ($\epsilon < -130\%$). This result indicates that the produced CO should exhibit highly ¹³C-depleted composition and may in turn explain the ¹³C-enrichment of early Martian CO₂. Under a reducing atmosphere, CO is likely transformed into simple organic compounds including aldehydes and carboxylic acids. Thus, if the scenario is correct, we predict that highly ¹³C-depleted organic matter may be deposited in early Martian sediments.

On early Earth, such an elevated level of CO may also explain the anomalous ¹³C-depletion in late Archean sediments from 2.7 to 2.5 Ga, which has been interpreted as the result of global methanotrophy [2]. Our recent study on sulfur photochemistry [3] also suggested that the late Archean atmosphere would have been very reducing, possibly including % level of CO or CH₄, in order to explain the S-MIF record including ³⁶S.

In summary, the atmosphere of the early Earth and Mars may have been rich in carbon monoxide, that could have played a significant role on the origin and evolution of biosphere.

[1] Schmidt et al. (2013) *PNAS* **110**, 17691-17696. [2] Hayes (1994) *Early Life on Earth* **84**, 220-236. [3] Endo et al. (2016) *EPSL* **453**, 9-22.