

Oxygen isotopic study of a CAI and an Al-rich Chondrule from Vigarano

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Calcium, Aluminum-rich inclusions (CAIs) are the the earliest forming solids of the Solar system that have been U-Pb dated to ~4568 Ma [1-3]. CAIs are suggested to form during the earliest epoch (~0-2 Ma) and some may have been (re)formed/ altered during the later high temperature events [4-8]. Oxygen isotopic studies of CAIs can therefore provide information about their formation and subsequent interaction with different oxygen isotopic reservoir in the early Solar System. A large igneous type B2 CAI (Vig CAI#1) and an Al-rich chondrule from Vigarano (CV~3.1-3.4) were selected for studying their oxygen isotopic composition to understand their petrogenetic history and compare with previous studied Magnesium isotopic records of the CAI. The Al-rich chondrule is an atypical object consisting of two large (~2cm) distinct olivine grains that sandwich Al-rich phases (plagioclase, glass, spinel). Multiple traverse to the core of CAI and a few analysis within the Al-rich chondrule were made in multi collection mode using faraday cups of Secondary ion mass spectrometer 1280 HR2 at Heidelberg University. Vigarano CAI #1 has uniform oxygen isotopic composition corresponding to $\Delta^{17}\text{O}$ of $\sim -1\pm 2$ permil within this large CAI of ~2 cm. The previously obtained $\delta^{25}\text{Mg}$ isotopic composition of the melilite also show the similar constant composition within the CAI. Al-rich chondrule on the other hand displays some variability amongst the olivine and Al-rich phases but has relative ^{16}O rich oxygen isotopic composition in the range of $\Delta^{17}\text{O}$ of $\sim -4\pm 2$ permil. Similar oxygen isotopic composition has been observed previously in CAIs and Al-rich chondrule. $^{26}\text{Al}/^{27}\text{Al}$ of $(4.89\pm 0.38)\times 10^{-5}$ of Vigarano CAI #1 implies that the CAI #1 formed at ~0.1 Ma in a “planetary like” ^{16}O poor oxygen isotopic environment.

- [1] MacPherson G. J. (2007) *Treatise on Geochemistry* (Heinrich, D.H., Karl, K.T. (Eds.), Pergamon, Oxford, 1-47. [2] Amelin et al. (2010) *EPSL* 300, 343-350. [3] Bouvier & Wadhwa (2010) *Nature geo.* 3, 637-641. [4] Jacobsen B. et al. (2008) *EPSL*, 272, 353-364. [5] Kita N. T. et al. (2012) *GCA*, 86, 37-51. [6] Mishra R. K. and Chaussidon M. (2014) *EPSL*, 390, 318-326. [7] MacPherson G. J. et al. (2012) *EPSL*, 331, 43-54. [8] Larsen K. K. et al. (2011) *APJL*, 735, L37.