Evidence of fossil record of ⁷**Be in a CAI: Implications for the early Solar system**

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⁷Be, that decays to ⁷Li with half life of 53.06 ± 0.12 days [1], is a key short-lived now-extinct radionuclide to derive information about early solar system event and processes. Lithium-berylium-boron (Li-Be-B) isotope systematics studies in the first forming solar system solids, Ca-Al-rich inclusions (CAIs) provide a unique opportunity of utilising two isotope decay systematics of ⁷Be and ¹⁰Be to ⁷Li and ¹⁰B, respectively, to understand cosmochemical/astrophysical conditions and plausibly also chronology of the events and processes in the early solar system [2-4]. A first unambigous detection of 7Be along with fossil records of 10Be corresponding to 7Be/9Be of (1.2±1.0)×10-3 (95% conf.) and ${}^{10}\text{Be}/{}^{9}\text{Be}$ of $(1.6\pm0.32)\times10^{-3}$ is being inferred from the regression of the in situ isotopic data obtained using secondary ion mass spectrometer in a pristine type B CAI from Efremovka (CV~3.1-3.4). Isotopic records of ⁷Be, ¹⁰Be and ²⁶Al in a type B CAI from Efremovka (E40) allow to make following very important inferences: (1) Nascent Sun underwent multiple episodes of enhanced magnetic activity (2) the later episode of enhanced irradiation occuring at the end of "class I" stage of pre-main sequence evolution was more intense (3) Irradiation is the prime source of ⁷Be and also ¹⁰Be. An intense irradiation by a super flare (X-ray luminosity $L_x \approx 10^{32}$ ergs) during the terminal class I stage of a CI (carbonaceous ivuna \approx solar) composition precursors near the reconnection region for about an year can concurrently explains the isotopic properties (7Be, 10B, 26Al), morphology (texture, modal grain sizes), and petrology (mineral compositions) of CAI, along with preservation of faster diffusing lithium isotope records.

[1] Jaeger, M., et al. *Phys. Rev. C* 54, 423-424 (1996). [2] McKeegan, K. D., et al. *Science* 289, 1334-1337 (2000). [3] Chaussidon, M., et al. *Geochim. Cosmochim. Acta* 70, 224-245 (2006). [4] Marhas, K. K., et al. *Science* 298, 2182-2185 (2002).