## THE NATURE AND TIMING OF REFRACTORY ELEMENT FRACTIONATION EVENTS IN THE EARLY SOLAR NEBULA

G. J. MACPHERSON<sup>1</sup>, A. N. KROT<sup>2</sup>, AND K. NAGASHIMA<sup>2</sup>

<sup>1</sup> Smithsonian Institution, Washington, D.C. 20560. USA (macphers@si.edu);

<sup>2</sup> Univ. of Hawai'i at Manoa, Honolulu, HI

Volatility-controlled elemental fractionation occurred in the pre-solar nebula at 4.56 Ga. The classic model [1] is equilibrium condensation of a hot solar-composition gas within a monotonically-cooling disk. Solid and melt evaporation contributed to further fractionation [2]. Evidence of these nebular processes is preserved in chondrules and calcium-aluminum-rich inclusions (CAIs; the first solid objects formed in the solar system) from chondrite meteorites.

 $^{26}$ Al (t<sub>1/2</sub>~ 0.7 my) existed in the early solar nabula, and it can serve as a high-precision relative chronometer. Histograms of initial <sup>26</sup>Al/<sup>27</sup>Al ratios derived from internal isochrons in CAIs are trimodal, with peaks at  $5 \times 10^{-5}$ .  $4.2 \times 10^{-5}$ , and 0. Including data from chondrules adds a fourth peak at 0.7×10-5. If CAI formation / reprocessing was a continuous process, no peaks would be expected. Discrete events must have occurred and, as CAIs formed very near the infant Sun [3, 4], events in the Sun itself likely were responsible. [5] proposed that major FU Orionis flare-ups in the Sun were the cause. Such flare-ups wax and wan very quickly (few years) at infrequent intervals and may be due to non-uniform accretion of disk material onto the protostar. The effects can be profound: a recent outburst in V883 Ori may have expanded its ice line (water/snow boundary) out to  $\sim 40$ AU [6]. A presumed similar effect on the rock line leads to the possibility of multiple condensation and re-condensation events in the earliest solar system, greatly affecting how we view elemental fractionation. Fractional, dis-equilibrium condensation is the more realistic model.

Our new measurements of initial  ${}^{26}\text{Al}/{}^{27}\text{Al}$  in fine-grained CAIs having Group II REE patterns (signatures of fractional condensation) indicate a major re-condensation event when  ${}^{26}\text{Al}/{}^{27}\text{Al}=5\times10^{-5}$ , requiring a prior event that conceivably formed those CAIs with little or no  ${}^{26}\text{Al}$ .

**References:** [1] Grossman L (1972) *GCA* **36**: 597–619. [2] Davis A. M. and Richter F. M. (2014) *In Treatise on Geochemistry* 2<sup>nd</sup> edition, 361-395; [3] McKeegan K. D. et al. (2011) *Science* **332**, 1528. [4] McKeegan K. D. et al. (2000) Science **289**: 1334–1337. [5] MacPherson G. J. (2017) 48th LPSC, Abst. #2719; [6] Cieza L. A. et al. (2016) *Nature* **535**, 258-261.