

## **U-corrected Pb-Pb ages of chondrules: what are we dating?**

CONNELLY, J.N.<sup>1</sup>, BOLLARD, J.<sup>1</sup> AND BIZZARRO, M.<sup>1</sup>

<sup>1</sup>StarPlan, University of Copenhagen, Denmark.

The disagreement between chondrule ages by different chronometers has been attributed to the typically older absolute Pb-Pb ages being unreliable due to the complex histories of chondrules [1, 2]. Chondrules will inherit initial Pb ( $Pb_i$ ) from their precursor material that is not significantly evolved from the primordial Pb isotopic composition. The dual decay of  $^{235}\text{U}$  and  $^{238}\text{U}$  to  $^{207}\text{Pb}$  and  $^{206}\text{Pb}$ , respectively, generates radiogenic Pb ( $Pb_r$ ) with a unique radiogenic  $^{207}\text{Pb}/^{206}\text{Pb}$  ratio for any interval of time. Complete melting will mix/equilibrate original  $Pb_i$  and any  $Pb_r$  formed by that time, with some Pb likely lost by evaporation (potentially with Pb isotopic fractionation). The sum of mixing and fractionation of existing Pb will result in a new  $Pb_i$  that will be incorporated into any phase(s) able to accommodate Pb (e.g. sulphide). The phase(s) incorporating U will accumulate new  $Pb_r$ . The re-setting of  $Pb_i$  in Pb-hosting phases may happen multiple times until the last closure and final lock of the isotopic composition of  $Pb_i$ . From this point in time forward,  $Pb_r$  ingrowth in U-bearing phases will again define a unique radiogenic  $^{207}\text{Pb}/^{206}\text{Pb}$  ratio formed since last melting. We find that U is hosted by the mesostasis, a phase readily linked to the chondrule's last melting. As such, the derived  $Pb_r$  age must date this event alone, without any record of the pre-history. Incomplete re-mixing of Pb during partial melting will result in a ternary mixture of Pb that will fail to define a statistically valid line and, therefore, fail to define an age. All combined, a successfully defined Pb array defines the  $Pb_r$  value related to the time of last melting, independent of a complex history prior to the last re-mixing of Pb. Given the rapid cooling of chondrules, all robust ages from different chronometers should yield the same age that reflects the last full melting event. We attribute the offset between the  $^{26}\text{Al}$ - $^{26}\text{Mg}$  ages and Pb-Pb ages to the inhomogeneous distribution of  $^{26}\text{Al}$  in the protoplanetary disk. Using our Pb-Pb ages and  $Pb_i$  estimates for 22 individual chondrules, we infer that nebular chondrules started forming at the same time that calcium aluminium inclusions formed ( $4567.30 \pm 0.16$  Ma, [3]) and continued to form for ca. 3.6 Myr [4]. Furthermore, we define a primary chondrule formation episode within the first million years of the Solar System followed by a period of reworking that lasted up to 3.6 Myr.

[1] Kita, N. et al. *MAPS*, 48, 1383. [2] Budde, et al. (2016) *PNAS*, 113, 2886. [3] Connelly, J.N. et al. (2012) *Science*, 338, 651. [4] Bollard J. et al. (2017) *Science Advances*, submitted.