

High-pressure serpentinization, deep H₂ and abiotic methanogenesis

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Serpentinization is known to be an important source of H₂. When carbon is available in serpentinizing systems, abiotic hydrocarbons may form and play a key role in geobiological processes on Earth and possibly beyond. A great international effort is currently made on the study of serpentinizing systems and the related abiotic hydrocarbons in shallow terrestrial environments within the biosphere, such as mid-ocean ridges and ophiolites. However, the P and T region at which serpentinization may occur extends to much greater depths reaching more than 100km under subduction zone thermal regimes. High-P serpentinization is widely documented by both natural observations and geophysical studies. Despite that, the role of this process on the generation of H₂ and abiotic hydrocarbons, most notably CH₄, is unknown. The recent discovery of natural high-P processes responsible for the generation of large amounts of H₂ and abiotic CH₄ in deep serpentinizing systems (Vitale Brovarone et al., 2017, Nature Communications) now calls for a reassessment of the role of high-P serpentinization as a reservoir of deep abiotic hydrocarbons. This presentation summarizes the field, microstructural and stable isotope results, as well as their implications and long-term perspectives. Our data suggest that CH₄ can form by different reactions sustained by deep, dry fluids that were not documented before in these settings, and that may lead to the precipitation of C particles characterized by signatures equivalent to diamonds. These mechanisms have the potential to also occur and sustain pre-biotic processes at lower pressure within the biosphere. It is emphasized that deep serpentinization has the potential to produce CH₄ fluxes comparable to those estimated at mid-ocean ridges. This achievement represents an important means to better understand the relationships between biotic and abiotic processes in the biosphere, as well as for the detection and exploration of similar processes on other bodies of the solar system.