

**Partitioning of Si and S between  
solid and liquid in the Fe-Si-S system  
up to 25 GPa with implications for  
the distribution of Si and S in a  
partially solidified core**

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Planetary cooling leads to solidification of any initially molten core. Both S and Si have been suggested to be present in planet's cores. We conducted high-pressure experiments to investigate the partitioning behavior of Si and S in the Fe-Si-S system up to 25 GPa. The starting material contains ~ 6 wt.% S and ~ 4 wt.% Si. The liquidus temperature in this study is about 50 °C lower than that in the Fe-S binary system for the same S concentration in the melt at 21 GPa. Similar to the binary system, S prefers to partition into liquid and its content in the liquid increases with decreasing temperature. However, Si solubility in the solid phase is almost independent of temperature, controlled by the bulk Si content. The Si content in the liquid phase increases with temperature, accompanied with the decrease in the S content. We fitted the partition coefficient ( $K_D^{Si}$ ) of Si between solid and liquid in correlation with experimental P, T and S concentration in liquid. At same pressure, the  $\log(K_D^{Si})$  is only linearly dependent on  $1/T$ . The higher temperature (lower  $1/T$ ), the higher  $\log(K_D^{Si})$ . With increase of pressure, the slope of the linear correlation between  $\log(K_D^{Si})$  and  $1/T$  will decrease. The Si partition coefficient is also strongly dependent of the S concentration in liquid. Our experimental results have directly applications to the distribution of Si and S between the solid inner core and the liquid outer core in small terrestrial planets such as Mars and Mercury. Understanding the pressure effect on the partitioning coefficients will allow us to further constrain the distribution of Si and S in the Earth's core.