## Constraints on the W abundance of upper mantle from xenolithic peridotites

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Well-resolved <sup>182</sup>W variations compared to terrestrial standards have been discovered in ancient to modern mantlederived rocks. Given that  $^{182}\mathrm{Hf}$  decays into  $^{182}\mathrm{W}$  in the first 50 Ma of Solar system hystory, variations in <sup>182</sup>W reflect Hf/W fractionation early in Earth's history. Tungsten behaves as an incompatible element during igneous processes and W concentrations of mantle-derived magmas should be higher than their source, yet W concentrations of some mantlederived rocks with <sup>182</sup>W anomalies are often unexpectedly high (up to a few ppm). This enrichment is difficult to explain by igneous processes alone, given estimates for the W concentration of the modern mantle of  $\sim 8 \text{ ppb}$  [1]. The W concentration of mantle reservoirs are largerly based on data from mid-ocean ridge basalts, ocean island basalts, and ophiolites [1-3]. To understand the behaviour of W during mantle melting, and better constrain the W concentration of the upper mantle, we report W concentrations for well characterized mantle xenoliths from Hannouba and Yangyuan (China), and Shavaryn (Mongolia). Hannouba and Yangyuan samples have W concentrations ranging from 10 to 80 ppb; Shavaryn samples range from 20 to 150 ppb. Combined with major and trace element data, W concentrations of some of these mantle xenoliths suggest W enrichments occurred through re-fertilization and/or metasomatic overprinting. Nevertheless, other xenoliths show trends of W vs. MgO, Al<sub>2</sub>O<sub>3</sub>, CaO and SiO<sub>2</sub>, for whole rocks, clinopyroxenes and spinels consistent with W depletion accompanying partial melt removal. This suggests that the best preserved and fertile mantle xenoliths can be used to constrain the W abundance of the upper mantle.

[1] Arevalo and McDonough (2008) EPSL 273, 656-665. [2] Ireland et al. (2009) GCA 73, 4517-4530. [3] Babechuk et al. (2010) GCA 74, 1448-1470. [4] O'Driscoll et al., (2015) J PETROL 56, 1797-1828.