## Magnesite stability at the reduced mantle conditions from the metal saturation depth to the lower mantle

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Magnesite is suggested to be a major oxidized carbon storage phase in the mantle due to its wide P-T range of stability [1]. However, the presence of magnesite in the Earth's interior depends highly on the redox conditions. Large part of the mantle is significantly reduced, with detectable amount of Fe<sup>0</sup> dispersed in the silicate rocks [2]. Therefore, subducted carbonates would interact with Fe<sup>0</sup>. However, the mechanism of this interaction remains controversial.

We investigated MgCO<sub>3</sub>-Fe<sup>0</sup> interaction at 6-145 GPa and 800-2600 K. To determine influence of water on the reaction, additional experiments were conducted in the hydromagnesite-Fe<sup>0</sup> system at 6 and 16 GPa. Redox reactions in both systems occur with the formation of (Mg,Fe)O, graphite/diamond, and Fe<sub>3</sub>C/Fe<sub>7</sub>C<sub>3</sub>. The fundamental difference was connected with the reaction mechanisms. In the "dry" system solid state diffusion of components was suggested to be the major rate-limiting process with the reaction rate constant (k) in the range of  $10^{-11} - 10^{-13}$  at 6 GPa and 1400 - 1600 K. In the water-bearing system, reactions were governed by solubility of components in the water fluids, and were controlled by the hydrogen fugacity. Water acts as a transporting agent for the reactants. Calculations of the kinetic parameters clearly detected the increase of k by two order of magnitude in comparison with the "dry" system.

The sluggish kinetics of the MgCO<sub>3</sub>–Fe<sup>0</sup> reaction established in our study suggests that carbonates could survive during subduction from metal saturation boundary near 250 km depth down to the transition zone and even to the core-mantle boundary. At the same time, in the watercontaining systems magnesite would be completely reduced to carbide or diamond.

[1] Fiquet et al. (2002). Am. Min., 87:1261-1265 [2] Frost., McCammon (2008). Annu. Rev. Earth Planet. Sci. Lett., 36: 389-420.