

## Dating metal-silicate fractionation among chondrite parent bodies

J.L. HELLMANN<sup>1\*</sup>, T.S. KRUIJER<sup>1,2</sup> AND T. KLEINE<sup>1</sup>

<sup>1</sup>University of Münster, Institut für Planetologie, Münster, Germany. \*jan.hellmann@uni-muenster.de

<sup>2</sup>Lawrence Livermore National Laboratory, Livermore, USA

The distinct metal-to-silicate ratios of ordinary chondrites (defining the subgroups H, L and LL) most likely reflect nebular metal-silicate fractionation just prior to chondrite parent body accretion. Until now, however, the timing of this metal-silicate fractionation event, and whether it was coeval to chondrule formation, has not been determined. To address these issues, we applied the short-lived  $^{182}\text{Hf}$ - $^{182}\text{W}$  chronometer to a suite of ordinary chondrites, with the ultimate goal of constraining the timing of metal-silicate fractionation among the H, L and LL ordinary chondrite subgroups. Owing to their distinct metal-to-silicate ratios, H, L and LL chondrites exhibit different bulk Hf/W, and so each group should have followed distinct Hf-W evolutionary paths after metal-silicate fractionation. To determine these evolution paths and hence, the time at which the distinct Hf/W ratios were established, we obtained Hf-W isochrons for thirteen H, L and LL chondrites of petrologic types 4-6.

For all samples, precise Hf-W isochrons were obtained, corresponding to ages of between  $\sim 3$  and  $\sim 11$  Ma after formation of CAIs. For most LL chondrites, metal typically plots below the isochron defined by corresponding silicate fractions, indicating a slightly earlier closure of the Hf-W system in the metal. Overall, our results demonstrate that type 6 chondrites have younger Hf-W ages than type 4 chondrites, implying slower cooling for chondrites of higher petrologic type and consistent with a concentrically layered structure of chondrite parent bodies, where more strongly metamorphosed samples are located at greater depth [1]. In a plot of initial  $^{182}\text{W}$  isotopic composition vs. time, H, L and LL chondrites define distinct evolution paths that intersect at  $\sim 2$ -3 Ma after CAIs. Thus, the distinct bulk Hf/W (LL>L>H) were established at this time. This time of metal-silicate fractionation is consistent with Al-Mg ages of chondrules of  $\sim 2$  Ma after CAIs [e.g. 2]. As chondrule formation likely was followed by rapid accretion to a parent body [e.g. 3], our results thus show that chondrule formation, metal-silicate fractionation and chondrite parent body accretion all were coeval at ca. 2-3 Ma after CAI formation.

**References:** [1] Trieloff M. et al. (2003) *Nature*, 422, 502–506. [2] Kita N.T. and Ushikubo T. (2012) *MAPS*, 47, 1108–1119. [3] Alexander C.M.O. et al. (2008) *Science*, 320, 1617–1619.