## Noble gas systematics of Martian meteorite Northwest Africa 10416

## WILLIAM S. CASSATA<sup>1</sup>, CONNOR HILTON<sup>1</sup>, LARS E. BORG<sup>1</sup>, CARL $AGEE^2$

<sup>1</sup> Lawrence Livermore National Laboratory,

Livermore, CA, USA

<sup>2</sup> University of New Mexico, Albuquerque, NM, USA

Northwest Africa (NWA) 10416 is an olivine phyric Martian basalt with a trace element composition similar to (LREE) depleted shergotittes [1]. It is the second Martian basalt recovered to date that has crystaline, igneous feldspar, the other being NWA 8159 [2]. To better understand the petrogenentic relationship between NWA 10416 and other Martian meteorites, here we report He, Ne, Ar, Kr, and Xe isotopic measurements, (U-Th)/He chronometry, and <sup>40</sup>Ar/<sup>39</sup>Ar chronometry.

Preliminary  ${}^{40}$ Ar/ ${}^{39}$ Ar data are indicative of a 400 – 500 Ma crystallization age. Comparable ages have been obtained from the chemically similar depleted Shergottites Dar al Gani (DaG) 476 [3] and Yamato 980459 [4]. High temperature extractions likely associated with gas released from large (up to 1 mm) olivine crystals yield ages that exceed 1 Ga, which suggests that excess Ar is present in olivine crystals or they are of xenocrystic origin. The trapped component identified by isochron analysis indicates that NWA 10416 equilibrated with a Martian interior noble gas reservoir or has subsequently acquired terrestrial noble gases due to Antarctic weathering. Additional  ${}^{40}$ Ar/ ${}^{39}$ Ar measurements on mineral separates are underway.

Cosmic ray exposure ages were calculated from <sup>3</sup>He, <sup>21</sup>Ne, and <sup>38</sup>Ar measurements using the production rate estimates of [5] and chemical composition determined by ICP-MS. All three systems yield exposure ages of  $1.05 \pm 0.15$  Ma. Like the <sup>40</sup>Ar/<sup>39</sup>Ar data, similar exposure ages have been obtained from DaG 476 [6] and Yamato 980459 [7]. (U-Th)/He ages obtained from three whole-rock fragments of NWA 10416 are concordant at ~20 Ma, consistent with >95% loss of He during the ejection event that occurred at ~1 Ma.

This work was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344.

[1] Herd et al. (2016) LPSC Abstract #2527. [2] Walton et al. (2016) LPSC Abstract #1639. [3] Borg et al. (2003) GCA 67, 3519-3536. [4] Shih et al. (2005) AMR 18, 46-65. [5] Eugster & Michel (1995) GCA 59, 177-199. [6] Park et al. (2003) LPSC Abstract #1213. [7] Nishiizumi & Hillegonds (2004) Antarctic Meteorites XXVIII, 60-61.