

Mn-Cr systematics of D'Orbigny revisited

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For almost a decade, the angrite D'Orbigny has been used as an important anchor for Mn-Cr chronometry. While the U-Pb age of D'Orbigny is now tightly constrained [1, 2], very few high-precision data for $(^{53}\text{Mn}/^{55}\text{Mn})_0$ have been reported since the study of Glavin et al. [3]. A major concern about the $(^{53}\text{Mn}/^{55}\text{Mn})_0$ reported in Glavin et al. [3] is the use of second-order fractionation correction assuming terrestrial $^{54}\text{Cr}/^{52}\text{Cr}$ ratio. It has been shown by Trinquier [4] that the $\epsilon^{54}\text{Cr}$ of angrite is anomalous ($\epsilon^{54}\text{Cr} = -0.36 \pm 0.07$), so the use of such fractionation correction will result in a positive shift of the $\epsilon^{53}\text{Cr}$ values. Furthermore, if the errors associated with the $^{54}\text{Cr}/^{52}\text{Cr}$ ratios have not been propagated to the $\epsilon^{53}\text{Cr}$ values, the stated uncertainty will be an underestimated value. Recently, McKibbin et al. [5] reported $(^{53}\text{Mn}/^{55}\text{Mn})_0$ of $(3.54 \pm 0.18) \times 10^{-6}$ from the analysis of olivine using SIMS. However, the large uncertainty lowers the resolution of the Mn-Cr ages, making it less useful for cosmochemical studies.

Here we report high-precision Cr isotope data of whole-rock and mineral separates physically and chemically separated from D'Orbigny. The Cr isotope ratios were measured using TRITON Plus following the method described in [6]. Our new data show that the $\epsilon^{54}\text{Cr}$ values of the mineral separates are not uniform, possibly due to spallation effect. The $\epsilon^{53}\text{Cr}$ values are correlated with the $^{55}\text{Mn}/^{52}\text{Cr}$ ratios, defining a preliminary isochron whose slope corresponds to a $(^{53}\text{Mn}/^{55}\text{Mn})_0$ of $(3.35 \pm 0.04) \times 10^{-6}$. We have also determined the noble gas composition of the sample to calculate its exposure age, which was then used to investigate the contribution of spallation reaction on the $\epsilon^{53,54}\text{Cr}$ signatures. Further chronological implications of these new results, together with additional data from the analysis currently in progress, will be presented.

[References] [1] Amelin (2008) *GCA* **72**, 221-232. [2] Brennecka et al. (2012) *PNAS* **109**, 9299-9303. [3] Glavin et al. (2004) *MAPS* **39**, 693-700. [4] Trinquier et al. (2007) *ApJ* **655**, 1179-1185. [5] McKibbin et al. (2015) *GCA* **157**, 13-27. [6] Yamakawa et al. (2009) *Anal. Chem.* **81**, 9787-9794.