

Chemical characteristics of irghizite: Implications for the nature of the projectile

SHIRAI N., AKHTER R., EBIHARA M.

¹ Department of Chemistry, Tokyo Metropolitan University, Hachioji, Tokyo 192-0397, Japan, shirai-naoki@tmu.ac.jp

Irghizites and zhamanshinites are tektite-like objects and impactites from the Zhamanshin impact crater, respectively. There are obvious chemical differences between irghizites and zhamanshinites. Irghizites have higher Co and Ni abundances than those of zhamanshinites [1]. These enrichments have been considered to be due to the contamination of extraterrestrial materials [e.g., 1-3]. Based on the chemical compositions, it is likely that chondrite and iron meteorites are candidates of the impactor materials on the Zhamanshin crater [2-4]. However, no clear conclusion regarding the identification of impactor has been made. In this study, we determined the chemical compositions of irghizite, specially focusing on siderophile element in order to identify the meteoritic component.

One irghizite sample was carefully ground in a clean agate mortars. Homogenized powder sample was taken for neutron irradiation for two different periods at Kyoto University Research Reactor Institute.

Thirty elements were determined in irghizite by using INAA. Our analytical results are in good consistent with those of previous studies [1-4]. Being consistent with the results of the previous studies [1-4], our Ni and Co values are higher than those of zhamanshinites [1,3] and similar to those of mafic materials. We consider a possibility that Cr, Co and Ni abundances for irghizite are controlled by the addition of chondrites or iron meteorites to crustal materials. In plot of Ni vs. Cr, both mixtures with chondrite or iron meteorite could produce Cr and Ni abundances of irghizite. For refined identification, platinum group elements should be determined.

[1] Taylor S. R. and McLennan S. M. (1979) *GCA*, 43, 1551-1565. [2] Palme H. et al. (1981) *GCA*, 45, 2417-2424. [3] Koeberl C. and Fredriksson K. (1986) *EPSL*, 78, 80-88. [4] Mizera J. et al. (2012) *JRNC*, 293, 359-376.