## Origin of large soluble molecules in Murchison chondrite.

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Organic matter in Murchison is made of thousands of compounds varying in mass up to 2000 Da [1]. Their size is in-between the molecules detected in space [2] and macromolecules in meteorites [3]. They have recorded signatures of complexification during the transition from the dense molecular clouds to the disk. We seek signature related to one stage of the meteoritic material evolution.

65 g of Murchison were macerated in Methanol and Toluene (1:2) for 1 week. Extracts were recovered after centrifugation and stored in glass tubes. Mass spectra were acquired with a Thermo LTQ Orbitrap XL coupled with an Electrospay ionization (ESI) source, in the $150-1000 \mathrm{~m} / \mathrm{z}$ range at resolving power $\mathrm{m} / \Delta \mathrm{m}=100000$.

Detected ions are in the 150-750 Da range. The average mass is $\sim 350 \mathrm{Da}$ and the average diversity is 5.1 ions per Da. Each mass detected bears at least NH , the maximal number of heteroatoms is 2 for N and 3 for O . The average $\mathrm{H} / \mathrm{C}$ is 1.6 regardless of the mass. There is no convergence toward macromolecular-like low saturation. We interpret the periodicity in mass as a repetition of stoichiometric patterns. $\mathrm{CH}_{2}, \mathrm{H}_{2}$ and $\mathrm{C}_{5} \mathrm{H}_{8}$ are the most frequent. Chain rearrangements, cuts and cycling signatures with limited loss of hydrogen are all together consistent with randomized additions of $\mathrm{CH}_{2}$ and $\mathrm{H}_{2}$ This can occur on grain surface [4]. PDR can provide an efficient C bonds cut mechanism [5].
[1]Schmitt-Kopplin P. et al. (2010) P.N.A.S., 107, 7 pp. 2763-8. [2]Caselli P. and Ceccarelli C. (2012) A.A.Rev., 20, 1 p. 56. [3]Sephton M. a (2002) Nat. Prod. Rep., 19, 3 pp. 292-311. [4]Belloche A. et al. (2014) Science (80-. )., 345, 6204 p. 15841587. [5]Alata I. et al. (2015) A.\&A., 123 pp. 1-9.

