

## Cosmogenic Effects on Cr Isotopic Composition of Iron Meteorites

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<sup>53</sup>Mn-<sup>53</sup>Cr short-lived chronometer ( $t_{1/2}=3.7$  Myr) is a powerful tool for high resolution chronological studies of the early Solar System events occurred in the very first 10 Ma. However, cosmogenic effects, induced by spallation and thermal/epithermal neutron capture processes, could hamper the correct determination of radiogenic contribution to <sup>53</sup>Cr, therefore influence the accuracy of <sup>53</sup>Mn-<sup>53</sup>Cr chronometry, especially for samples with high Fe/Cr, Ni/Cr ratios and long exposure ages ([1,2]). Metal and olivine phases in meteorites that are often used for the determination of Mn-Cr age, have high Fe and/or Ni contents and could be vulnerable to cosmogenic effects. Thus it is necessary to examine the cosmogenic Cr isotopic signature in meteorites, and to find a proper method to correct for such effects on Cr isotopic composition in order to obtain meaningful Mn-Cr ages.

Iron meteorites are ideal samples for investigating cosmogenic effects on Cr isotopic composition because of their high Fe/Cr, Ni/Cr ratios and relatively long exposure ages. In this study, we analyzed 25 iron meteorite samples from 9 chemical groups using a Thermo Finnigan Triton multi-collector thermal ionization mass spectrometer at DTM. The samples display large coupled variations in  $\epsilon^{53}\text{Cr}$  and  $\epsilon^{54}\text{Cr}$  (up to  $\sim 250\epsilon$  for  $\epsilon^{53}\text{Cr}$  and up to  $\sim 1000\epsilon$  for  $\epsilon^{54}\text{Cr}$ ) with a ratio of approximately 1:3.9. This ratio is similar to that determined through three pieces of Carbo meteorites in [3] and is consistent with our modeling results ( $\sim 1:3.4$ ). We found that this correlation is very robust and is independent of the chemical composition (different Fe/Ni ratio) of the meteorites. Therefore, we can use this correlation to correct for cosmogenic effects on  $\epsilon^{53}\text{Cr}$ , because the pre-exposure  $\epsilon^{54}\text{Cr}$  (anomalies) are often known and the magnitude is small compared to the variation generated by cosmogenic effects.

[1] Shima, M. and M. Honda (1966) *EPSL*, **1**(2): 65-74. [2] Birck, J. and C. Allegre (1985) *Isotopic Ratios in the Solar System*, **1**: 21-25. [3] Qin L. et al. (2010) *GCA*, **74**(3): 1122-1145.