

Mineralogy of Eolian Sands at Gale Crater, Mars

C.N. ACHILLES¹, R.T. DOWNS¹, D.T. VANIMAN²,
A.S. YEN³, D.F. BLAKE⁴, R.V. MORRIS⁵, D.W.
MING⁵, E.B. RAMPE⁶, S.M. MORRISON¹, T.F.
BRISTOW⁴,
S.J. CHIPERA⁷, B.L. EHLMANN⁸, M.G.A.
LAPOTRE⁸,
K.S. EDGETT⁹, R. GELLERT¹⁰, A.H. TREIMAN¹¹,
K.V. FENDRICH¹, P.C. SARRAZIN¹², P.I. CRAIG¹¹,
J.A. CRISP³, J.P. GROTZINGER⁸, D.J. DES
MARAIS⁴,
J.D. FARMER¹³, J.M. MOROOKIAN³

¹Dept. of Geosciences, Univ. of Arizona, Tucson, AZ
85721, USA, cheriea@email.arizona.edu

²Planetary Science Institute, Tucson, AZ, USA

³Jet Propulsion Laboratory/Caltech, Pasadena, CA,
USA

⁴NASA Ames, Moffett Field, CA, USA

⁵NASA Johnson Space Center, Houston TX, USA

⁶Aerodyne Industries at NASA JSC, Houston, TX,
USA

⁷Chesapeake Energy, Oklahoma City, OK, USA

⁸California Institute of Technology, Pasadena, CA,
USA

⁹Malin Space Science Systems, San Diego, CA, USA

¹⁰Dept. of Physics, Univ. of Guelph, Guelph, ON,
Canada

¹¹Lunar and Planetary Institute, Houston, TX

¹²SETI, Mountain View, CA, USA

¹³Arizona State University, Tempe, AZ, USA

The Mars Science Laboratory rover, Curiosity, is using a comprehensive scientific payload to explore rocks and soils in Gale Crater, Mars [1]. Two eolian sands have been sampled during MSL's mission: the Rocknest sand shadow consisting of armored sediments [2] and Gobabeb, a sample from the active Bagnold dune field at the base of Mount Sharp. The CheMin X-ray diffraction instrument performed quantitative mineralogical analyses of the <150 μm portion of these sediments. Both deposits are dominated by basaltic minerals along with a significant portion of X-ray amorphous material. In addition, the mineral chemistry of olivine and plagioclase is unchanged based on modeled peak positions in the Rocknest and Gobabeb diffraction data. Bulk chemical analyses of the Gobabeb >150 μm fraction from APXS suggest that the mineralogy is more mafic than the CheMin analyzed fraction. This presentation will compare the active, Gobabeb, and inactive, Rocknest, sand deposits and use the mineralogy to explore the weathering histories and sources for each sediment.

[1] Grotzinger, *Space Science Reviews* (2012) **170** 5-56. [2] Blake *et al.* (2013) *Science* **341** 1239505.