

Evidence for change in C/O ratio during evolution of enstatite chondrites

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Unequilibrated Enstatite Chondrites (EH3&EL3) evolved from a source region with high C/O ratio (> 0.83). The nature of their birth place and mode of formation is controversially debated i. e., either as solar nebula condensates (SCON), or by impact melting (IM) in parental asteroids. These contrasting mechanisms would imply different mineral inventories, cosmochemical signatures, and would leave isotopic signatures undisturbed (in SCON) or alternatively homogenized (IM). We report mineralogical, cosmochemical and nanoSIMS isotopic results from primitive EH3s and the EL3s MS-17 and MS-177 (TC₃ asteroid) fragments. EH3 chondrites contain a menagerie of sulfides consisting of FeS, niningerite (Mg Fe Mn)S, djerfisherite, sphalerite, minor oldhamite (CaS), FeNi metal and graphite. Growth zoning in (Mg Fe Mn)S to contacting FeS reveals a dichotomy of trends: (1) normal (FeS decrease) & (2) reverse (FeS increase) latter attesting thermal events in different asteroids. ¹²⁹I/¹²⁹Xe dating of djerfisherite in the most primitive EH3 ALHA77295 indicates an age comparable to oldest age of chondrules in E-chondrites (**4564.2 ± 1.1 M.a**; King et al., 2013). Mineral inventory in EL3, in contrast consists of FeS, alabandite (Mn Fe Mg)S, sphalerite oldhamite, enstatite, FeNi metal and idiomorphic sinoite. Growth sequence of earliest condensates is oldhamite-sinoite-graphite thus indicating increase in C/O ratio after sinoite. REE inventory of CaS in EH3 depicts positive Eu and Yb anomalies. The pattern in EL3 is in contrast, flat with a negative Eu anomaly suggesting a different source region. NanoSIMS study of graphite various morphologies indicates clear dichotomy in δ¹³C and δ¹⁵N signatures in the same EL3 chondrite. Our results clearly negate the (IM) origin for both EH3 and EL3.