

Planetary scale Ca-isotope heterogeneity and the formation history of the Earth-Moon system

M. SCHILLER^{1*}, V. A. FERNANDES² AND M. BIZZARRO¹

¹Centre for Star and Planet Formation, Natural History Museum of Denmark, University of Copenhagen, Copenhagen, Denmark (*email: schiller@snm.ku.dk)

²Museum für Naturkunde, Leibniz-Institut für Evolutions- und Biodiversitätsforschung, Berlin 10115, Germany,

The Moon is thought to be the result of a collision between a Mars-sized object, Theia, and the Earth [1]. In this model, the Moon is dominantly composed of impactor material. The isotopic similarity in many elements between the Earth and the Moon requires that isotopic re-equilibration occurred after the impact through a vapour plume [2] or, alternatively, Theia was nearly isotopically identical to Earth [3]. Distinguishing between these scenarios require investigating the isotope composition of highly refractory elements, as these have long isotopic equilibration timescales.

As one of the most refractory elements, Ca is ideally suited for this task. Moreover, it was recently shown that the abundance of ⁴³Ca, ⁴⁶Ca and ⁴⁸Ca is variable and correlated amongst inner solar system primitive and differentiated parent asteroids [4]. Using improved techniques, we have analyzed the mass-independent ⁴³Ca/⁴⁴Ca and ⁴⁸Ca/⁴⁴Ca compositions of several lunar and Martian meteorites by HR-MC-ICPMS. Our preliminary data suggest the existence of Ca-isotope heterogeneity between the Earth, Moon and Mars. The ⁴³Ca/⁴⁴Ca and ⁴⁸Ca/⁴⁴Ca compositions of Martian meteorites are depleted by -1.1 ± 0.9 and -21.2 ± 5.6 ppm relative to Earth's mantle. In contrast, lunar meteorites are enriched ⁴³Ca/⁴⁴Ca by $+1.5 \pm 1.0$ ppm but have a terrestrial ⁴⁸Ca/⁴⁴Ca composition ($+0.4 \pm 3.9$ ppm). The observed isotopic variability cannot be explained by neutron capture during exposure. This suggests that, although their isotopic make-up is similar, the Earth, Mars and Theia sampled isotopically distinct precursor material, arguing against the proposal of a uniform inner solar system Ca isotope reservoir [3]. Moreover, preliminary Ca-isotope data for two enstatite chondrites are not consistent with this material being the precursor of Theia. Thus, the observed planetary and asteroidal Ca-isotope heterogeneity cannot be solely attributed to admixing of ⁴⁸Ca-rich outer system material to a depleted inner solar system composition.

[1] W. K. Hartmann and D. R. Davis (1975) *Icarus* **24**, 504-515. [2] K. Pahlevan and D. J. Stevenson (2007) *EPSL* **262**, 438-449. [3] N. Dauphas et al. (2014) *EPSL* **407**, 96-108. [4] Schiller et al. (2015) *GCA* **74**, 4844-4864.