

## **U-Pb zircon ages and Sr-Nd isotopic composition of Neoproterozoic magmatism, Dionisio-Sierra de los Ríos block, NE Uruguay**

E. PEEL\*<sup>1</sup>, R. MUZIO<sup>1</sup> AND M. A. S. BASEI<sup>2</sup>

<sup>1</sup>Instituto de Ciencias Geológicas, Facultad de Ciencias, Universidad de la República. Iguá 4225 CP 11400, Montevideo, Uruguay (\*correspondence: elena@fcien.edu.uy)

<sup>2</sup>Instituto de Geociências, Universidade de São Paulo. Rua do Lago 562, Cidade Universitaria, São Paulo, Brazil.

The Dionisio-Sierra de los Ríos block consists of various plutonic rocks of Neoproterozoic age, all related to the collision between Río de la Plata and Kalahari cratons [1 and references therein]. This block belongs to the Aiguá-Pelotas Batholith of the Dom Feliciano Belt, also known as the granitic belt [2]. Their composition range from granites, granodiorites and sienogranites, showing variable degree of deformation from isotropic to milonitic textures. Geochemically they are high K calc-alkaline rocks, showing magmatic arc to post collisional signature.

A total of five samples presenting igneous zircons with overprinted late-post-magmatic features were analyzed. The obtained U-Pb zircon ages (LA-ICP-MS) yielded a range from 625 to 600 Ma for each sample, implying 25 Ma time interval for their emplacement. The isotopic data obtained in the same samples indicate  $\epsilon_{\text{Nd}}$  values range from -4.2 to -6.4, while  $^{87}\text{Sr}/^{86}\text{Sr}$  range varies from 0.70817 to 0.71759.

The ages obtained for these rocks are coherent with the ones indicated for other granitic bodies belonging to the Aigua and Pelotas batholiths [3][4], developed to southwest and northeast respectively. The ages obtained here are interpreted as indicative of the initial magma generation (~625 Ma) and emplacement (~600 Ma). The isotopic and petrologic data suggest crustal participation in their genesis.

[1] Sánchez-Bettucci et al. (2010) *INT GEOL REV*, **52**: 1, 51-78. [2] Basei et al. (2000) *Tectonic evolution of South America, 31<sup>st</sup> IGC*, pp. 311-334. [3] Oyhantçabal et al. (2010) *INT J EARTH SCI*, **100(2-3)**: 379-390. [4] Philip & Machado (2005) *J S AM EARTH SCI*, **19**: 461-478.