

Elastic wave velocity of antigorite up to 5 GPa and 500°C

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Introduction

In some previous studies on the elastic wave velocities of serpentine (e.g., [1]), it has been reported that the ratio of compressional wave velocity to shear wave velocity (V_p/V_s) for serpentine is anomalously high compared to other major minerals in the Earth's upper mantle. As the results, seismological studies often discuss that the anomalous V_p/V_s near the plate boundary would be the evidence for the presence of serpentine around subduction zones (e.g., [2]). However, the elastic properties of serpentine have not been fully understood yet because P-T conditions in the previous experimental studies were limited to relatively low P-T (e.g., [3]) or high-P but room T [4]. It is, therefore, important to determine the elastic wave velocities of serpentine in wide range of P-T conditions for better understanding of H₂O transportation to the Earth's interior around subduction zones.

Experimental method

High P-T experiments were conducted using a Paris-Edinburgh cell at Beamline 16 BM-B, HPCAT of the Advanced Photon Source. An Al₂O₃ buffer rod was used for transmitting elastic waves between WC anvil and sample. A LiNbO₃ transducer, which generates and receives both V_p and V_s , was attached to the top WC anvil. 30 MHz and 20 MHz electrical sine waves were used to determine V_p and V_s , respectively. The details on ultrasonic measurement setup are described in Kono *et al.* [5]. Powdered gem-quality antigorite was used as a starting material in order to obtain isotropic elastic data. The measurements were done up to about 5 GPa and 500°C.

Results and discussion

The obtained V_p for antigorite increases monotonically with increasing P, whereas V_s is almost constant with P. Both V_p and V_s decrease slightly with increasing T. The detailed data and geophysical implication will be presented.

[1] Christensen (1996) *JGR* **101**, 3139-3156. [2] Kamiya & Kobayashi (2000) *GRL* **27**, 819– 822. [3] Watanabe *et al.* (2007) *Earth Planets Space* **59**, 233-244. [4] Bezacier *et al.* (2013) *JGR* **118**, 527–535. [5] Kono *et al.* (2012) *Rev. Sci. Instrum.* **83**, 033905–033908.