

Collisional history of asteroid Itokawa

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In situ extraterrestrial samples returned for study (e.g., the Moon) are crucial in understanding the origin and evolution of the Solar System as, contrary to meteorites, they provide a known geological context for the samples and their analyses [1]. Asteroid 25143 Itokawa is a rubble pile asteroid consisting of reaccumulated fragments from a catastrophically disrupted monolithic parent asteroid, and from which regolith dust particles have been recovered by the Hayabusa space probe [1].

We analysed two dust particles using Electron Backscatter Diffraction (EBSD) and ⁴⁰Ar/³⁹Ar dating techniques, the latter using an ARGUS VI mass spectrometer. One of the grains showing signs of 15-25 GPa impact shock pressure, yielded a ⁴⁰Ar/³⁹Ar plateau age of 2.3 ± 0.1 Ga. These data represents the first precise ⁴⁰Ar/³⁹Ar plateau age ever obtained from a single-grain dust particle recovered in situ from an extra-terrestrial body. Therefore, our approach opens a new era of single-grain analyses for precious, rare and small particles of extraterrestrial material.

Our diffusion and porosity-P-T models show that the relatively low pressure and high temperature involved in the impact process can be reconciled only if the asteroid was already made of porous material at ~ 2.3 Ga and thus, if asteroid Itokawa was already formed, thereby providing a minimum age for catastrophic asteroid breakup [1] [2]. A second particle shows no sign of deformation indicating shock pressure of < 10 GPa and a calculated maximum temperature of $\sim 200^\circ$. This low temperature estimate is supported by an apparent ⁴⁰Ar/³⁹Ar age of 5.4 ± 1.6 Ga (i.e. an age of 3.8 - 4.6 Ga) indicating a lack of isotopic resetting for this particle. This suggests that the breakup of Itokawa's parent was a relatively low-temperature process and occurred on a pre-shattered parent body [3].

[1] Abe et al., *Science* (2006). [2] Tsuchiyama, *Elements* (2014). [3] Michel et al., *Icarus* (2004).