

Ru stable isotope analyses of meteorites by double spike MC-ICPMS

T. HOPP¹, M. FISCHER-GÖDDE¹ AND T. KLEINE¹

¹Institut für Planetologie, University of Münster, 48149 Münster, Germany (timo.hopp@wwu.de)

The abundances of highly siderophile elements (HSE) in the primitive upper mantle of the Earth are higher than those expected for equilibrium metal-silicate partitioning during core formation in Earth (e.g. [1]). This HSE excess is commonly explained by the addition of a late veneer of primitive material to the Earth's mantle after core formation was complete [e.g. 2]. However, non-chondritic relative abundances of Pd and Ru in the Earth's mantle indicate that either the late veneer was comprised of material distinct from known groups of chondrites (e.g. [1] [3]) or represents a mixture of chondritic and differentiated meteoritic material [4]. Alternatively, the mantle abundances of some HSE are at least partly influenced by metal-silicate equilibration at high temperatures and pressures [e.g. 5]. For instance, the concentrations of Au, Pt and Pd can be explained by high PT core formation (e.g. [5-7]). To distinguish between the addition of a late veneer and high-PT metal-silicate partitioning during core formation as the dominant processes establishing HSE abundances, the investigation of stable isotope fractionations in HSE are useful. This is because metal-silicate equilibration of HSE should impart a strong isotope fractionation in the silicate, due to the very strong affinity of HSE for metal.

We have developed a new ⁹⁸Ru-¹⁰¹Ru double spike for the precise measurement of small mass-dependent Ru isotope variations in meteorites and terrestrial samples. The isotope measurements were conducted using the Neptune *Plus* MC-ICPMS at the University of Münster, and the data reduction is made off-line using the double spike toolbox [8]. Our first preliminary results show that carbonaceous and ordinary chondrites as well as IIIAB iron meteorites are isotopically fractionated compared to an Alfa Aesar Ru standard solution. There also is a hint of small Ru stable isotope variations among the investigated meteorites, but these differences are small and not yet clearly resolved and need further investigation through the analyses of a more comprehensive set of samples.

[1] Becker et al (2006) *GCA* **70**, 4528-4550. [2] Kimura et al (1974) *GCA* **38**, 683-701. [3] Walker (2009) *Chem. Erde* **69**, 101-125. [4] Fischer-Gödde and Becker (2012) *GCA* **77**, 135-156. [5] Righter et al (2008) *Nat. Geosci.* **1**, 321-424. [6] Danielson et al (2005) *LPSC XXXVI*, 1955. [7] Cottrell et al (2006) *GCA* **70**, 1565-1580. [8] Rudge et al (2009) *Chem. Geol.* **265**, 420-431.