

## Is SO<sub>2</sub> photolysis the source for Archean mass-independent sulfur isotope signatures?

ANDREW R. WHITEHILL<sup>1,\*</sup> AND SHUHEI ONO

<sup>1</sup>Department of Earth, Atmospheric, and Planetary Sciences, Massachusetts Institute of Technology, Cambridge, MA

\*Correspondence to: arwhite@mit.edu

Sulfur isotope mass-independent fractionation (S-MIF) is observed in Archean rocks and is currently considered the best evidence for a low-oxygen early atmosphere [1]. Results from photochemical models suggest that the production and preservation of S-MIF places strict constraints on atmospheric oxygen levels during the Archean [2]. Much of our interpretation of S-MIF, however, is based on the hypothesis that SO<sub>2</sub> photolysis by ~200 nm radiation is the source of these anomalous isotope signatures.

We studied the sulfur isotope systematics of the photolysis of SO<sub>2</sub> under a range of experimental conditions, studying the effects of temperature (-60 to 25°C), total pressure (250 mbar to 1 bar), SO<sub>2</sub> pressure (down to 0.1 mbar) and the addition of other gases (CH<sub>4</sub>, O<sub>2</sub>). The systematics of the isotope effects observed during experiments support a major contribution from isotopologue self-shielding to S-MIF production during SO<sub>2</sub> photolysis [3].

If SO<sub>2</sub> photolysis were the dominant source of the Archean S-MIF signatures, experimental results suggest:

- (1) SO<sub>2</sub> photolysis is occurring under high column densities of SO<sub>2</sub> (i.e. self-shielding regime), and thus S-MIF is predominantly produced in localized, transient volcanic plumes rather than in the background atmosphere.
- (2) SO<sub>2</sub> photolysis is occurring at low total (i.e. N<sub>2</sub>) pressures. Either atmospheric pressure during the Archean was lower than 1 bar, or photolysis must have been occurring at altitudes above 10 km (assuming a 1 bar atmosphere)
- (3) Large mass-dependent fractionations are required prior to eventual preservation in the rock record to reduce the large δ<sup>34</sup>S values produced during photolysis.
- (4) High SO<sub>2</sub>:H<sub>2</sub>S ratios of volcanic gases are required for elemental sulfur to contain sufficiently high S-MIF signatures (i.e. contribution from SO<sub>2</sub> relative to H<sub>2</sub>S)

We conclude that the isotope systematics of SO<sub>2</sub> photolysis [3] are hard to reconcile with the geological data by itself. CS<sub>2</sub>, SO and S<sub>2</sub> also undergo photolysis by predissociation and could be alternative source reactions for the Archean S-MIF signatures.

[1] Farquhar *et al* (2000) *Science* **289**, 756 [2] Pavlov and Kasting (2002) *Astrobiology* **2**(1), 27 [3] Ono *et al* (2013) *J. Geophys. Res. Atmos.* **118**, 2444