

Organic Carbon inventory of the Tissint meteorite.

A. STEELE¹, F. MCCUBBIN², L.G. BENNING³,
S. SILJESTRÖM⁴, G. CODY¹, Y. GOREVA⁵, E. HAURI⁶,
J. WANG⁶, A. KILCOYNE⁷, M. GRADY⁸, C. SMITH¹¹,
C. FREISSINET¹², D. GLAVIN¹², A. BURTON¹³, M. FRIES¹⁴,
J. BLANCO³, M. GLAMOCLJA¹, K. ROGERS¹,
S. MIKHAIL⁶ AND J. DWORKIN¹²

¹University College London, Gower St, London.
(stealie@mac.com.)

²Inst. of Meteoritics, Dept. of Earth and Planetary Sci, Univ.
New Mexico, Albuquerque, New Mexico, 87131 USA

³School of Earth and Environment, University of Leeds,
Leeds, LS2 9JT, UK.

⁴Department of Chemistry and Materials, SP Technical
Research Institute of Sweden, 501 15 Borås Sweden.

⁵Department of Mineral Sciences, Smithsonian Institution,
Washington, DC. 20013-7012 USA.

⁶Department of Terrestrial Magnetism, Carnegie Institution of
Washington, 5241 Broad Branch Rd, Washington DC.
20015 USA.

⁷Advanced Light Source, 1 Cyclotron Road, MS 7R0222,
LBNL, Berkeley, California 94720. USA.

⁸Centre for Earth, Planetary, Space and Astronomical
Research. Open University, Milton Keynes, Walton Hall,
Milton Keynes, MK7 6AA, UK.

⁹Université de Toulouse, UPS-OMP, IRAP, Toulouse, France.

¹⁰CNRS, IRAP, 9 Av. colonel Roche, BP 44346, 31028,
Toulouse Cedex 4, France.

¹¹Department of Mineralogy, The Natural History Museum,
Cromwell Road, London, SW7 5BD. U.K.

¹²NASA Goddard Space Flight Center, Greenbelt Road,
Maryland 20771, USA.

¹³Catholic University of America, NASA Goddard Space
Flight Center, 8800 Greenbelt Road, Greenbelt, Maryland,
20771, USA.

¹⁴Planetary Science Institute, 1700 East Fort Lowell, Suite
106, Tucson, Arizona, 85719 USA

The fall of the Tissint meteorite has provided a unique opportunity to study a minimally contaminated piece of Mars. Martian organic carbon has been detected previously in igneous basalts and the carbonates of ALH 84001. Analysis of sealed maskelynite inclusions using *in situ* techniques including Raman, NanoSIMS, ToFSIMS, STXM and TEM, coupled with whole rock analysis by stepped combustion, GCMS and evolved gas analysis has revealed an inventory of organic compounds containing -CH, -CN, -CNO, -COOH, -CO and aromatic complexes. These are spatially resolved to known inorganic catalysts, i.e. magnetite, pyrite, nickel containing pyrrhotite and clays. Furthermore there is a release of nitrogen containing organics above 600°C, at which temperature $\delta^{15}\text{N}$ is $\sim +40\text{‰}$. These results show that Mars has an inventory of organic carbon and nitrogen containing molecules that are probably produced through abiological hydrothermal activity.

PTX properties of FeCl₂-bearing fluids at elevated PT conditions

MATTHEW STEELE-MACINNIS^{1*}, PILAR LECUMBERRI-SANCHEZ¹ AND ROBERT J. BODNAR¹

¹Department of Geosciences, Virginia Tech, Blacksburg VA
24061, USA (*correspondence: mjmaci@vt.edu)

Iron chloride is a significant component of saline aqueous fluids in many ore-forming environments, commonly occurring at concentrations of up to several mass percent [1]. Thus the effect of FeCl₂ on the phase equilibria of aqueous fluids is a significant factor in interpreting fluid evolution in iron-rich systems such as tin-tungsten deposits and some porphyry copper deposits. However, there are few available experimental data on the phase equilibria and thermometric properties of iron-bearing aqueous fluids [2].

We have used the synthetic fluid inclusion technique to investigate the pressure-temperature conditions along the locus of critical points in the system H₂O-FeCl₂, from 0 to 35 wt% FeCl₂. Our results show that the effect of FeCl₂ on the phase equilibria of aqueous fluids is unlike that of other divalent cation chlorides such as CaCl₂ or MgCl₂. Specifically, the locus of critical points for H₂O-FeCl₂ fluids occurs at significantly lower pressure (at a given temperature) compared to the critical curve of the system H₂O-NaCl. For example, at 500 °C the critical point for H₂O-FeCl₂ fluids occurs at 480 bar, compared to about 560 bar for H₂O-NaCl. The low pressure along the H₂O-FeCl₂ locus of critical points implies that immiscibility (boiling) of FeCl₂-rich fluids can occur only at relatively low pressures, or relatively shallow levels in the crust. These results allow us to reinterpret the conditions of mineralization at boiling, Fe-rich hydrothermal systems.

[1] Yardley (2005) *Econ. Geol.* **100**, 613-632. [2] Liebscher (2007) *Rev. Mineral. Geochem.* **65**, 15-47.