

Airborne measurements of atmospheric trace gases via infra-red laser absorption spectroscopy

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Recent laboratory and field studies demonstrated that replacing lead-chalcogenide tuneable diode lasers by cw operated quantum cascade lasers (QCL) results in sensitivity improvements of mid-IR TDLAS systems by a factor 2 to 3. Therefore, the MPI-C three laser TRacer In-situ Tdlas for Atmospheric Research (TRISTAR) was equipped with 3 QCL emitting at 1268.98, 2158.30, and 1759.72 cm⁻¹ to measure CH₄, CO and HCHO, respectively. In October 2005 the modified TRISTAR instrument was installed on a Lear Jet 35A as part of a scientific payload to study the photochemistry over the tropical rainforest in South America during the GABRIEL campaign. A second deployment was during fall 2006 and summer 2007 as part of the HOOVER campaign to study HO_x and its precursors in the upper troposphere over Europe. Since 2012 the instrument has been successfully flown on the new HALO aircraft during the TACTS/ESMVAL and OMO-Europe missions. These missions investigated the influence of convection and stratosphere-troposphere-transport on the photochemistry of the tropopause region. The performance of the instrument during these airborne campaigns was examined for the three species and precisions for CO and CH₄ were measured in the field to be 0.5% and 0.8% respectively (2σ). The 1σ detection limit for HCHO was ~500 pptv for a 2 second average, while post-flight signal averaging over a 2 minute time interval resulted in a 150 pptv detection limit.

Phase diagrams of FeO and Fe-Si alloys

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Earth's core is less dense than pure iron, implying the presence of one or more lighter element(s) [1] such as Si, O, or S [2]. The phase diagrams of iron alloyed with these elements at high pressures and temperatures (*P-T*) are critical input for understanding the thermodynamics of these systems. Here we present results on FeO and a suite of Fe-Si alloys.

High *P-T* conditions (up to 200 GPa) were generated using a laser-heated diamond anvil cell. In situ X-ray diffraction to determine crystal structures was performed at beamline 13-ID-D of the Advanced Photon Source, Argonne National Laboratory. Melting was determined from diffuse X-ray scattering, by laser power-temperature relationships, and by temperature-emissivity relationships.

We have determined the melting curve of FeO [3] and clarified the location and slope of the B1/B8 phase transition [4]. We also identified an insulator-metal transition [5]. The B1 metallic phase of FeO is the stable phase at conditions of Earth's lower mantle and outer core, with possible implications for the high *P-T* character of Fe-O bonds, magnetic field propagation, and lower mantle conductivity.

FeSi has the B20 structure at 1 bar, the B2 structure at high pressures, and a wide two-phase field in between [6]. Fe-9Si has the hcp structure at high *P* and low *T*, and converts to an hcp+B2 mixture and then fcc+B2 with increasing temperature [6]. Fe-16Si has the DO₃ structure at low pressures and is an hcp+B2 mixture at higher pressures [7]. We have also measured melting temperatures for each alloy. Phase diagrams in *P-X* and *T-X* space imply that the stable phase of Fe-Si alloy at inner core conditions for compositions that match the observed density deficit is an hcp+B2 mixture [6].

[1] Birch (1952), *J. Geophys. Res.* **57**, 227-286. [2] Allègre *et al* (1995), *Earth Planet. Sci. Lett.* **134**, 515-526. [3] Fischer and Campbell (2010), *Am. Mineral.* **95**, 1473-1477. [4] Fischer *et al* (2011a), *Earth Planet. Sci. Lett.* **304**, 496-502. [5] Fischer *et al* (2011b), *Geophys. Res. Lett.* **38**, L24301. [6] Fischer *et al* (in review). [7] Fischer *et al* (2012), *Earth Planet. Sci. Lett.* **357-358**, 268-276.