

Helium and neon isotopes as mantle tracers

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Noble gas isotopes, especially He isotopes, are widely used tracers for the formation and evolution of the Earth's mantle and atmosphere. Basic concepts on mantle structure and evolution are primarily based on the interpretation of mantle ³He as reflecting primordial, undegassed mantle material. In general, the relative enrichment of ³He observed in MORBs and OIBs, compared to atmospheric values, is interpreted in terms of the retention of primordial He by the mantle throughout the history of the Earth. Primordial He isotopic ratios in terrestrial matter are largely thought to be solar-like, with the deviation from those solar-like ratios increasing during Earth's history caused by the production of radiogenic ⁴He. Thus high ³He/⁴He ratios are interpreted to represent deep mantle material, whereas ratios around 8 R_A (R_A stands for the atmospheric ³He/⁴He ratio of 1.39 x 10⁻⁶) are thought to be representative for the upper mantle. Based on He, Ne and Ar fusion data of fresh, submarine volcanic glasses of a number of Mid-Atlantic Ridge off-axis seamounts we show that melt formation and evolution can have a larger impact on He than on e.g. Ne resulting in a decoupled behavior of He from other elements, such as e.g. Ne or Pb. All obtained He data are indistinguishable from the MORB range. In contrast, Ne isotopic compositions are much more primitive than MORB. Combined He, Ne and Ar systematics show that the source region of these seamounts experienced a preferential loss of He compared to Ne and Ar. This He loss, combined with subsequent ⁴He production, resulted in the decoupling of the He isotope systematics from Ne and Pb. Thus, among He and Ne only Ne has preserved the evidence that a primitive mantle component contributed to the formation of the investigated seamounts. As these seamounts are not fed from a mantle plume being derived from the deep mantle, the primitive Ne component resides within the upper mantle, implying that primitive noble gases are not necessarily indicative for deep mantle material. Our studies point out the necessity of obtaining Ne data in addition to He for the modeling of mantle formation and evolution and correct source characterization.

He-Ne-Ar isotope constraints on the nature and origin of high ³He/⁴He mantle

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Olivine phenocrysts from early Tertiary picrites from Baffin Island (BI) and West Greenland have ³He/⁴He = 38-50 R_A (n=24). The high ³He/⁴He are consistent with derivation from a mantle reservoir that is relatively undegassed compared to the depleted upper (MORB-source) mantle. Although Ne and Ar concentrations are typically 2 orders of magnitude lower than in basaltic glasses a low blank crusher is allowing Ne and Ar isotope determinations of high-³He/⁴He olivines. On a conventional 3-isotope plot, the BI picrites are indistinguishable from the Iceland and solar trends; the highest ²⁰Ne/²²Ne is 11.3. ⁴⁰Ar/³⁶Ar are typically less than 1,000. ³⁸Ar/³⁶Ar are indistinguishable from air values providing no evidence for solar Ar in the high-³He/⁴He mantle. Most samples define a trend in ²⁰Ne/²²Ne-⁴⁰Ar/³⁶Ar space that is consistent with mantle end-member with ⁴⁰Ar/³⁶Ar of 6,000-8,000. ⁴He*/²¹Ne* and air-corrected ³He/²²Ne imply the magmatic noble gases have suffered intense fractionation. This is supported by co-variation of ⁴He*/⁴⁰Ar* and ³He/³⁶Ar. The elemental fractionation is consistent with recent magmatic degassing and provides no evidence for an ancient degassing event necessary if the high-³He/⁴He mantle was a residue of early Earth depletion. The BI picrites plot on a trend in ⁴He*/²¹Ne*-³He/²²Ne space defined by basaltic glasses from Iceland. This is distinct from samples of Kola intrusives and we tentatively propose different degassing/depletion histories for the high-³He/⁴He mantle domains.