

## 6.3.26

### Disturbed $^{26}\text{Al}$ - $^{26}\text{Mg}$ system among CAIs and chondrules in mildly metamorphosed chondrites

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The  $^{26}\text{Al}$ - $^{26}\text{Mg}$  chronometer is useful to determine the relative formation time between Ca, Al-rich Inclusions (CAIs) and chondrules in unequilibrated chondrites. Previous studies have shown that CAIs formed within a few 0.1 million years (Myr) and chondrules formed 1-3 Myr after CAIs, which are also referred to as the time scale of the proto-planetary disk. As the initial  $^{26}\text{Al}/^{27}\text{Al}$  ratio in the solar system is very small ( $5 \times 10^{-5}$ ), plagioclase with high  $^{27}\text{Al}/^{24}\text{Mg}$  ratios ( $>100$ ) is often used in order to detect small  $^{26}\text{Mg}$  excess. However, there are numbers of CAIs and chondrules without any detectable  $^{26}\text{Mg}$  excess, indicating young formation ages of more than 4 Myr.

In this study, we present our recent  $^{26}\text{Al}$ - $^{26}\text{Mg}$  data on CAIs and chondrules in mildly metamorphosed chondrites (Ningqiang carbonaceous chondrite and Y82038 H3.2 chondrite). A Type C CAI in Ningqiang (W2#3) showed detectable  $^{26}\text{Mg}$  excess ( $\sim 1\%$ ) in Ca-pyroxene with a high referred  $^{26}\text{Al}/^{27}\text{Al}$  initial ratio ( $\sim 5 \times 10^{-5}$ ) as other CAIs. However, there is no detectable  $^{26}\text{Mg}$  excess ( $< 2\%$ ) in plagioclase, corresponding to the age of  $> 5\text{Myr}$ . A plagioclase-olivine inclusion (W2#4) consists of a dendritic and a poikilitic texture. The plagioclase in the poikilitic area showed correlated  $^{26}\text{Mg}$  excesses, giving the  $^{26}\text{Al}$  age of  $\sim 2.5\text{Myr}$ . The dendritic area showed higher  $^{27}\text{Al}/^{24}\text{Mg}$  ratios without any detectable  $^{26}\text{Mg}$  excess, which plot off the above isochron. These data indicate that the isotopic disturbance occur  $> 3.7\text{Myr}$ . The glass in type IA chondrule in Y82038 (F2) showed a limited  $^{26}\text{Mg}$  excesses ( $\sim 1\%$ ) that do not correlate with the  $^{27}\text{Al}/^{24}\text{Mg}$  ratios and give an upper limit of 3.6Myr. The  $^{27}\text{Al}/^{24}\text{Mg}$  ratios in the glass range 90-140, which is extremely higher than those in typical type IA in LL3.0-3.1 ( $< 10$ ).

These results indicate that the young  $^{26}\text{Al}$  ages of CAIs and chondrules are probably caused by mild metamorphism on their parent bodies 4-5 Myr after the CAI formation. There is a tendency of Mg removal from plagioclase or glass during the metamorphism, that caused apparent young ages on the  $^{26}\text{Al}$ - $^{26}\text{Mg}$  isochron diagram.

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### Contemporaneous formation of chondrules and CAIs inferred from the $^{26}\text{Al}$ - $^{26}\text{Mg}$ clock

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We report high-precision Mg isotope measurements of CAIs and chondrules from the Allende (CV3) and Dar al Gani 313 (LL3) meteorites determined by MC-ICP-MS. Analysed material comprised ca. 0.2-0.5 mg whole-rock fragments of chondrules or CAIs sampled with boron carbide micro-drills (300  $\mu\text{m}$ ) from polished slabs of the meteorites. Following table-top digestion in HF-HNO<sub>3</sub>, Mg was purified by cation exchange chromatography.  $^{26}\text{Mg}$  excesses are reported as  $\delta^{26}\text{Mg}^*$ , which is the per mil deviation of internally normalised (exponential law)  $^{26}\text{Mg}/^{24}\text{Mg}$  values from the average internally normalised  $^{26}\text{Mg}/^{24}\text{Mg}$  of two bracketing standards (DSM3), with each sample being analysed at least three times. The reproducibility of the  $\delta^{26}\text{Mg}^*$  value estimated from the BCR-2 basaltic standard is 0.023‰ (2 sd; n = 10). Al/Mg was determined by MC-ICP-MS to  $< \pm 2\%$ .

$(^{26}\text{Al}/^{27}\text{Al})_0$  values inferred from the CAI dataset range from  $3.5 \pm 1.0 \times 10^{-5}$  to  $6.3 \pm 1.8 \times 10^{-5}$ , indicating that these objects formed over a period of 0.6 Ma, or that some CAIs have been reworked or re-equilibrated with the host. An internal Al-Mg isochron (n = 5) for one CAI yielded  $(^{26}\text{Al}/^{27}\text{Al})_0 = 5.3 \pm 0.5 \times 10^{-5}$ , identical to the canonical  $(^{26}\text{Al}/^{27}\text{Al})_0$  value of  $5.0 \times 10^{-5}$ . Chondrules were selected for micro-drilling on the basis of their Al/Mg ratios, using a laser ablation system interfaced to an MC-ICP-MS to quickly screen a large number of chondrules. Seven ferromagnesian chondrules ( $^{27}\text{Al}/^{24}\text{Mg} = 0.117-0.274$ ) and two aluminium-rich chondrules ( $^{27}\text{Al}/^{24}\text{Mg} > 0.75$ ) were selected from Allende, while five ferromagnesian chondrules from Dar al Gani 313 with  $^{27}\text{Al}/^{24}\text{Mg} = 0.035$  to 0.145 were measured. Excess  $^{26}\text{Mg}$  was detected in 4 ferromagnesian and two aluminium-rich chondrules from Allende, while no measurable excess was detected in chondrules from Dar al Gani 313. The  $(^{26}\text{Al}/^{27}\text{Al})_0$  values inferred from the Allende chondrules range from  $1.7 \pm 0.6 \times 10^{-5}$  to  $4.8 \pm 0.6 \times 10^{-5}$ , suggesting that these objects formed over a period of ca. 1 Ma, and that some chondrules are as old as the oldest CAIs.