3.5.P11

Noble gas composition of the Réunion hotspot: Inferences for global mantle noble gas systematics

$\underline{J. \ HOPP} \ \text{AND} \ M. \ TRIELOFF$

Mineralogisches Institut, Universität Heidelberg, Im Neuenheimer Feld 236, D-69120 Heidelberg (jhopp@min.uni-heidelberg.de)

The Réunion mantle plume source is geochemically characterized by a very homogeneous source composition in He and Ne. The ${}^{3}\text{He}{}^{4}\text{He}\text{-ratios}$ of 13 ± 0.5 Ra (1 Ra = atmospheric ratio) are in between the values found for the MORB-reservoir (8±1 Ra) and for more primitive hot spots like Loihi (Hawaii) (up to 40 Ra). This intermediate composition is also reflected in the (air corrected) ${}^{21}\text{Ne}{}^{22}\text{Ne}$ ratio [1,2,3] corresponding to an intermediate contribution of nucleogenic ${}^{21}\text{Ne}$. Thus, Réunion is an attractive locality for noble gas studies: (1) the homogeneity points to a well mixed mantle reservoir (single component system) and (2) the intermediate character can serve as a connecting link between the composition of Loihi-like plume sources and the MORB-reservoir, that any successful global model regarding mantle noble gases has to account for.

The isotopic composition of Ar and especially Xe is less well constrained. In this study we reanalyse samples, which previously showed a high contribution of mantle Ne [2]. A relatively high precision in Ne isotope composition is aspired as a precondition to discriminate for atmospheric components in Ar and Xe. First results give a mantle 40 Ar/ 36 Ar ratio of about 10,000 (extrapolated to a mantle 20 Ne/ 22 Ne ratio of 12.5). Xe isotopes follow the mantle trend [3], with 129 Xe/ 130 Xe and 136 Xe/ 130 Xe values up to 6.93 and 2.36, respectively. In the light of our results, we will discuss the implications for current models in mantle noble gas geochemistry.

References

- [1] Hanyu T. et al., (2001) EPSL 193, 83-98.
- [2] Staudacher Th., Sarda P., and Allègre C.J. (1990) Chem. Geol. 89, 1-17.
- [3] Trieloff M., Kunz J., and Allègre C.J., (2002) EPSL 200, 297-313.