

Noble Gas Investigations of Early Proterozoic Lamproites from the Eastern Baltic Shield

Thomas Wiersberg (wiers@gfz-potsdam.de)¹, Samuel Niedermann (nied@gfz-potsdam.de)¹,
Jörg Erzinger (erz@gfz-potsdam.de)¹, Lev K. Levsky (lev@ad.igpp.ras.spb.ru)² &
Kirill I. Lokhov (kirill@kl1407.spb.edu)

¹ GeoForschungsZentrum Potsdam, Department 4.2, Telegrafenberg, 14473 Potsdam, Germany

² Institute for Precambrian Geology and Geochronology, Russian Academy of Sciences, Makarova emb. 2, RUS-199034 St. Petersburg, Russia

We present results of noble gas abundances and isotopic data of early Proterozoic lamproites (whole rock and mineral separates) from the eastern Baltic Shield (Kostamuksha/Karelia craton).

Ar isotope data from K-rich mineral separates (mica and kfs), released by stepwise heating, define a linear correlation when plotted in a $^{40}\text{Ar}/^{36}\text{Ar}$ versus $^{40}\text{K}/^{36}\text{Ar}$ diagram (Fig. 1). The slope of the regression line corresponds to an age of 1223 ± 13 Ma which agrees well with Sm-Nd and Rb-Sr dating (1235 Ma, Belyatskii et al, 1997). The Ar isotopic data of K-free minerals (cpx, quartz, carbonate) reveals the presence of excess argon which was also found in a fluid inclusion component, released by crushing of whole rocks. Diffusion of in-situ produced Ar from the matrix into the fluid inclusions is not likely, because the $^{40}\text{Ar}/^{36}\text{Ar}$ ratios in the fluid inclusions are not related to those in the crystal matrix. Therefore, we regard interaction of the rock material with a external fluid phase as a more likely explanation. The Ne isotopic ratios of crushed samples, when plotted in the Ne three-isotope diagram, display mixing between air and a crustal component (Kennedy et al., 1990). Therefore, any possible mantle component seems to be swamped by a fluid component with crustal origin.

In order to determine the $^{40}\text{Ar}/^{36}\text{Ar}$ ratio of the trapped component, we used a plot (not shown) of $^{21}\text{Ne}/^{22}\text{Ne}$ versus $^{40}\text{Ar}/^{36}\text{Ar}$ of crushed whole rock samples and melted K- and F-free mineral separates. Because of the different rare gas abundance ratios in the end members, the mixing line shows a more hyperbolic than linear array, indicating an upper limit for the $^{21}\text{Ne}/^{22}\text{Ne}$ ratio of 0.046. This corresponds to a $^{40}\text{Ar}/^{36}\text{Ar}$ ratio of ~ 5500 for the trapped component. Although a similar ratio is postulated for the Kola plume source (Marty et al., 1998), we do not suggest a mantle origin of the fluid for our samples, because their $^{20}\text{Ne}/^{22}\text{Ne}$ ratios are lower than atmospheric.

The neon stepwise heating data of whole rocks and micas show the presence of nucleogenic ^{21}Ne and ^{22}Ne , produced by Wetherill reactions $^{18}\text{O}(\alpha,n)^{21}\text{Ne}$ and $^{19}\text{F}(\alpha,n)^{22}\text{Ne}$ (β^+) ^{22}Ne , in proportions distinct from the crustal composition observed

in fluid inclusions. To assess this quantitatively, fluorine determinations are in progress.

The U/Th-He method yields ages of 100 - 450 Ma, indicating > 90% diffusion loss of He for most of the samples. No evidence of excess ^3He , either by crushing or by stepwise heating, was found. Measured $^3\text{He}/^4\text{He}$ ratios of melted samples are in good agreement with calculated $^3\text{He}/^4\text{He}$ production ratios via (n, α)-reaction of ^6Li (Mamyrin and Tolstikhin, 1984). $^3\text{He}/^4\text{He}$ ratios of crushed samples display the crustal ratio.

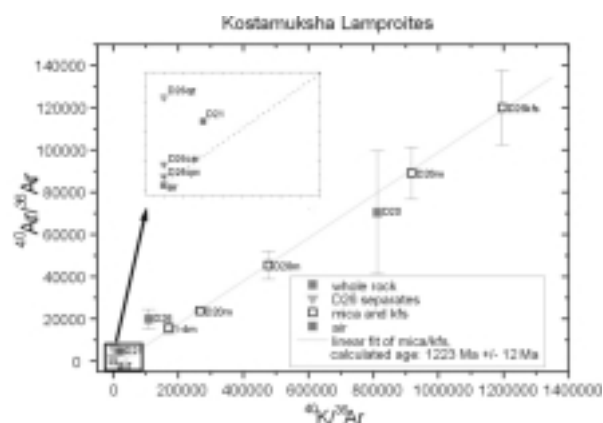


Figure 1: $^{40}\text{Ar}/^{36}\text{Ar}$ vs. $^{40}\text{K}/^{36}\text{Ar}$ plot of Kostamuksha lamproites, showing the presence of excess argon for K-free minerals. The slope of the regression line derived from K-rich minerals corresponds to an age of 1223 ± 13 Ma.

Belyatskii BV, Nikitana LP, Savaa EV, & Levsky LK, *Geochem. Int.*, **35**, 575-579, (1997).

Kennedy BM, Hiyagon H, & Reynolds JH, *Earth Planet. Sci. Lett.*, **98**, 277-286, (1990).

Mamyrin BA, & Tolstikhin I, *Helium Isotopes in Nature. Elsevier*, (1984).

Marty B, Tolstikhin I, Kamensky I L, Nivin V, Balaganskaya E, & Zimmermann J-L, *Earth Planet. Sci. Lett.*, **164**, 179-192, (1998).